



Slab-derived lithium isotopic signatures in mantle xenoliths from northeastern North China Craton

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ABSTRACT

In-situ lithium (Li) concentrations and $\delta^7\text{Li}$ of olivine, orthopyroxene (opx) and clinopyroxene (cpx) in mantle xenoliths from the Quaternary Longgang maar provide insights into the nature and evolution of the lithospheric mantle beneath the easterly part of the North China Craton. SIMS analyses show Li partitioning of olivine $>$ opx \geq cpx and Li isotopic disequilibrium in each sample. Olivines in harzburgite are generally homogeneous, with Li abundances in the range of normal mantle, but their $\delta^7\text{Li}$ (-0.9 to -4.0) are lower than normal mantle olivine $\delta^7\text{Li}$ values of $\sim +4 \pm 2$. Olivines in cpx-rich lherzolite and pyroxenite have higher Li contents and normal mantle-like $\delta^7\text{Li}$ values. The olivine $\delta^7\text{Li}$ values in the harzburgite and cpx-poor lherzolites are much lower than those published for worldwide peridotites, showing the peculiarity of these peridotites. The Li abundances and $\delta^7\text{Li}$ in the olivines show a positive correlation, indicating mixing between a low- $\delta^7\text{Li}$ melt and normal mantle. Collectively, Li elemental and isotopic compositions of the xenoliths suggest that low $\delta^7\text{Li}$ in olivines reflect a slab-derived Li isotopic signature in the lithospheric mantle. These observations, coupled with geochemical and geophysical evidence, suggest that the lithospheric mantle has been modified three times by (1) low- $\delta^7\text{Li}$ melt likely derived from the subducted Pacific plate (leading to the low $\delta^7\text{Li}$ of olivines), (2) high-degree partial melting (decreasing Li abundances of harzburgite and cpx-poor lherzolites close to the normal mantle), and (3) refertilization by melts mainly from the asthenosphere (producing the normal mantle-like $\delta^7\text{Li}$ and high Li contents in cpx-rich lherzolite and pyroxenite). The melts/fluids derived from the subducted Pacific plate could play a crucial role in the destruction of the craton.

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1. Introduction

The lithospheric mantle beneath the North China Craton (NCC) was old, thick, cold and typical of an Archean craton in the Paleozoic, but it had become younger, thinner, hotter and more “oceanic” by the Cenozoic. This is shown by the contrasting compositions of mantle xenoliths entrained in the Paleozoic kimberlites and Cenozoic basalts on the craton (e.g., Fan and Menzies, 1992; Fan et al., 2000a; Griffin et al., 1998; Menzies and Xu, 1998; Menzies et al., 1993, 2007; Rudnick et al., 2004; Xu, 2001; Zheng et al., 2001). Xenoliths from the Ordovician diamondiferous kimberlites in Mengyin and Fuxian (Fig. 1) have Archean Os model ages and Re depletion ages (Gao et al., 2002; Zhang et al., 2008a), providing strong evidence that at least some portions of the lithospheric mantle beneath the NCC were Archean and that the lithosphere was thick (~ 200 km) at the time of volcanism. In contrast, geophysical evidence (Chen et al., 2006; Yuan, 1996) and

geochemical data of mantle xenoliths in Cenozoic basalts suggest that the present lithosphere is much thinner (60–120 km) (Fan et al., 2000a; Gao et al., 2002; Griffin et al., 1998; Menzies and Xu, 1998; Menzies et al., 1993; Xu, 2001). The dramatic physicochemical changes of the lithospheric mantle suggest destruction of the NCC during the Phanerozoic time (e.g., Fan and Menzies, 1992; Menzies et al., 1993; O'Reilly et al., 2001; Xu, 2001; Zhang et al., 2002, 2003) or decratonization (Yang et al., 2008) due to the loss of typical characteristics of ancient cratons. As such the NCC offers an excellent opportunity to address the process and mechanism of continental lithospheric evolution.

It is generally accepted that the destruction of the NCC is evident in the large-scale thinning of the lithosphere, the increase in on-craton volcanism and the change in the compositions of the mantle xenoliths, e.g., from highly refractory harzburgites to fertile lherzolites and pyroxenites in mineral compositions and from high ϵ_{Sr} and low ϵ_{Nd} to low ϵ_{Sr} and high ϵ_{Nd} in isotopic compositions (e.g., Menzies et al., 2007; Zhang et al., 2009a and references therein). The character of the lithospheric mantle beneath the NCC also changed from typical cratonic into oceanic-type. However, the transformation process of the craton is still not well understood.

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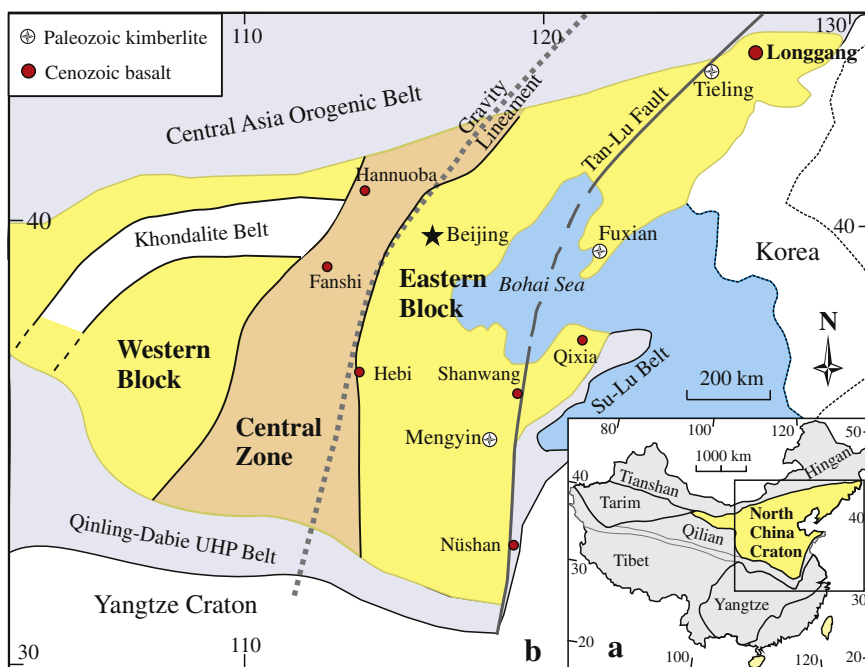


Fig. 1. Simplified geological map showing the localities of Paleozoic kimberlites and xenolith-bearing Cenozoic basalts mentioned in the text. The tectonic subdivisions of the North China Craton are from Zhao et al. (2008). Tan–Lu fault zone and south–north gravity lineament are two major linear zones. Inset shows location of the craton relative to other blocks and fold belts.

Our objective is to investigate the nature of the lithospheric mantle beneath the northeastern NCC in terms of Li isotopes and to use this information to further constrain the process by which the lithospheric mantle was transformed. Towards that end, we have analyzed in situ Li abundances and isotopic compositions of mantle xenoliths from Longgang region, which is located in the northeastern part of the NCC. Petrology, whole-rock chemistry, Sr–Nd and Os isotopic compositions of Longgang peridotite xenoliths have been documented previously (Fan et al., 2000b; Wu et al., 2003; Xu et al., 2003).

2. Geological background

The NCC is one of the oldest continental nuclei in the world, containing crustal rocks as old as 3.8 Ga (Liu et al., 1992). The basement of the NCC formed by the amalgamation of the Eastern Block and Western Block along the Central Zone (Fig. 1) in the Paleoproterozoic at ~1.85 Ga (Zhao et al., 2000, 2008). From its formation to the Mesozoic the NCC remained quiescent with respect to magmatic and tectonic events, except for the eruption of kimberlites in the middle Ordovician (Lu et al., 1998). Since the Late Mesozoic this craton has experienced widespread magmatism, as manifested by the emplacement of voluminous Late Mesozoic granites, mafic intrusions and volcanic rocks (Yang et al., 2003; Zhang et al., 2004) and extensive Cenozoic basalts carrying abundant mantle xenoliths (Tang et al., 2006; Zhou and Armstrong, 1982). These events and large-scale lithospheric thinning suggest reactivation of the NCC.

Based on petrological, geochemical and seismological data, the destruction of the NCC mainly occurred in the Eastern Block where the significantly thinned lithosphere is 60–100 km thick, while the Western Block was relatively stable and the lithosphere beneath this block is about 200 km thick (e.g., Chen et al., 2006; Fan et al., 2000a; Griffin et al., 1998; Menzies et al., 1993, 2007 and references therein). Maximum thinning of the lithosphere in the Eastern Block occurred along the Tan–Lu Fault and its neighboring regions where the thickness of lithosphere in this region is less than 75–80 km. Backarc extension caused by westward subduction of Pacific plate since the Cretaceous

may have played a crucial role in lithospheric thinning (Zhu and Zheng, 2009).

The Longgang volcanic field is located in Jingyu and Huinan counties of Jilin Province, China and contains more than 160 Quaternary volcanoes with 6 maars. These volcanoes form a volcanic cluster covering an area of about 2000 km² of the northeastern part of the NCC (Fig. 1). The earliest eruption of the Longgang volcanoes took place at about 2 Ma, and youngest eruption only about 1600 years B.P. Many of the volcanoes erupted in the Early Pleistocene (0.6 Ma). Most of the volcanoes have cinder cones composed of lavas and pyroclastic deposits such as bombs, block, cinders and ash.

3. Sample description

Mantle xenoliths in Longgang volcanic field are very abundant and large with diameters ranging mostly between 5 and 20 cm. Spinel-facies lherzolite is the most common rock type with minor spinel harzburgite, pyroxenite and composite spinel lherzolite xenolith interlayered with pyroxenite veins. This is similar to the case of Hannuoba basaltic field in the north margin of the central NCC (Chen et al., 2001; Fan and Hooper, 1989).

We selected five representative mantle xenoliths collected from the Longgang maar, which erupted at about 1 Ma. These samples are very fresh without any secondary alteration minerals. Primary minerals in the xenoliths are olivine, orthopyroxene (opx), clinopyroxene (cpx) and spinel. Samples 06JY02, 06JY06 and 06JY31 are spinel facies lherzolites with coarse-grained and porphyroclastic textures (Fig. 2). Most opx grains in the lherzolites are very coarse (2–6 mm in diameter) with a maximum diameter of over 1 cm. A large opx porphyroclast in 06JY06 shows exsolution lamellae. Samples 06JY02 and 06JY06 are relatively poor in cpx with similar modal mineral contents (olivine 64–64%, opx 25–26%, cpx 6–7%, spinel 3–5%) and are referred to as “cpx-poor” lherzolites in this study. In contrast, sample 06JY31 is cpx-rich lherzolite (olivine 60%, opx 18%, cpx 19%, spinel 3%) with pyroxenite–spinel clusters. Clinopyroxenes in the samples show a spongy texture. Sample 06JY30 is a pyroxenite vein interlayered in a

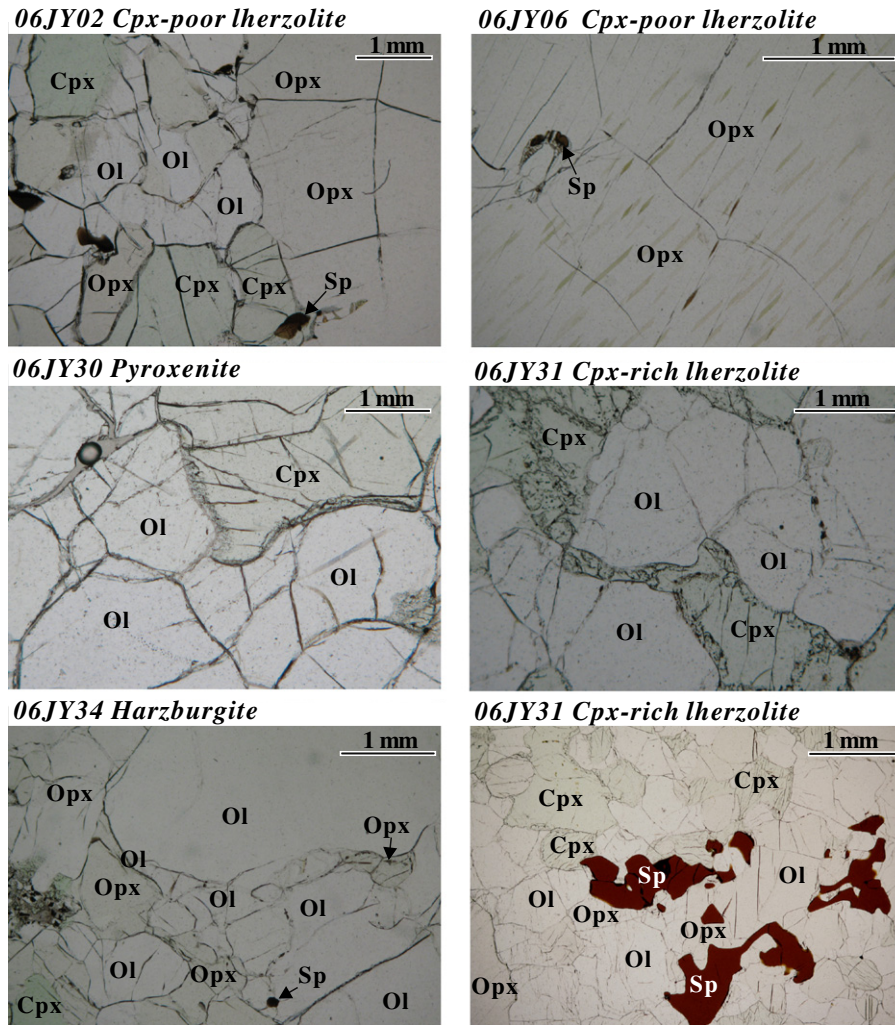


Fig. 2. Photomicrographs of Longgang spinel-facies xenoliths. Lherzolite 06JY02 has a coarse-grained, porphyroclastic texture with cpx breakdown. Lherzolite 06JY06 shows exsolution lamellae in a large opx porphyroblast. Clinopyroxene in the pyroxenite 06JY30 has a spongy texture. Cpx-rich lherzolite 06JY31 show clusters of coexisting cpx, opx and spinel and the spongy texture of cpx. Harzburgite 06JY34 shows inequigranular texture and breakdown features. A large olivine contains spinel inclusions. Cpx are interstitial along the grain boundaries of other minerals.

cpx-rich lherzolite and has modal mineral contents of olivine 30%, cpx 63%, opx 4%, and spinel 3%.

Sample 06JY34 is a coarse-grained spinel harzburgite (olivine 74%, opx 20%, cpx 3%, spinel 3%) with an inequigranular texture and mineral breakdown features (Fig. 2). Coarse grains of olivine and opx are up to 6 mm in diameter. Large olivines are kinked and some of them contain small spinel inclusions. Cpx grains (<1 mm) are smaller than the coexisting olivine and opx. A pyroxene–spinel cluster was observed in this sample.

In general, the distribution of opx and cpx is heterogeneous in these peridotite xenoliths. Clinopyroxenes in lherzolite 06JY06 and harzburgite 06JY34 are generally interstitial found along the grain boundaries and/or triple junctions. There are no hydrous phases (e.g., amphibole and phlogopite) in these xenoliths.

4. Analytical methods

Major element compositions of xenolith minerals (Table 1) were obtained with a JEOL JXA8100 electron probe microanalyzer (EPMA) at the Institute of Geology and Geophysics, Chinese Academy of Sciences, which was operated at an accelerating voltage of 15 kV and 10 nA beam current, 5 μm beam spot and 10–30 s counting time on peak. The precisions of all analyzed elements are better than 1.5%. Natural (jadeite [NaAlSi₃O₆] for Na, Al and Si, rhodonite

[MnSiO₃] for Mn, sanidine [KAlSi₃O₈] for K, garnet [Fe₃Al₂Si₃O₁₂] for Fe, Cr-diopside [(Mg, Cr) CaSi₂O₆] for Ca, olivine [(Mg, Fe)₂SiO₄] for Mg) and synthetic (rutile for Ti, 99.7% Cr₂O₃ for Cr, Ni₂Si for Ni) minerals were used for standard calibration.

Li concentrations and isotopic ratios were measured on gold-coated thin sections using Cameca IMS-1280 ion microprobe at the Institute of Geology and Geophysics, Chinese Academy of Sciences, following the procedures described in Zhang et al. (2010). A 13 kV, 10–20 nA O⁺ primary beam was focused on an area of 20 μm in diameter spot of the minerals. By sputtering the sample with the primary beam, the positive secondary ions produced were accelerated through 10 kV and measured at medium mass resolution ($M/\Delta M \sim 1100$), with a transmitted field of 125 μm, an aperture field of 4000, and an energy window of 60 eV without energy offset. The primary beam position, entrance slits, contrast aperture, magnetic field and energy offset were automatically centered before each measurement. Secondary ions were counted on mono-collection pulse counting mode. Forty cycles were measured with counting times of 12, 4 and 4 s for ⁶Li, background at the 6.5 mass and ⁷Li, respectively. The counting rate for ⁷Li ranges from 30,000 to 100,000 cps, depending on the Li content of the sample and primary beam intensity. A 60-s presputtering without raster was applied before analysis. Cpx samples BZ226 and BZ.CG, Ol sample BZ29 and Opx sample BZ226 (Decitre et al., 2002) were used as standards to correct the measured

Table 1
Major elemental compositions (wt.%) of minerals in Longgang mantle xenoliths.

Sample	Mineral	SiO ₂	TiO ₂	Al ₂ O ₃	Cr ₂ O ₃	FeO	MnO	MgO	CaO	Na ₂ O	K ₂ O	NiO	Total	Mg#
06JY02	Ol	40.4	0.00	0.02	0.01	9.35	0.18	48.8	0.10	0.03	0.04	0.35	99.3	90.4
	Opx	54.5	0.08	4.39	0.39	6.05	0.14	32.5	0.87	0.16	0.01	0.10	99.2	90.6
	Cpx	51.4	0.32	5.74	0.72	2.97	0.09	15.8	20.8	1.35	0.04	0.04	99.2	90.5
	Sp	0.04	0.11	52.7	12.2	11.0	0.11	20.3	0.01	0.03	0.02	0.32	96.9	76.9
06JY06	Ol	40.0	0.01	0.02	0.04	9.86	0.15	48.4	0.09	0.01	0.01	0.33	98.9	89.8
	Opx	54.7	0.11	3.65	0.50	6.28	0.14	32.4	0.84	0.09	0.01	0.09	98.8	90.3
	Cpx	51.9	0.38	4.86	0.87	3.16	0.08	16.0	21.8	0.98	0.00	0.05	100.1	90.1
	Sp	0.01	0.24	44.8	19.7	13.3	0.15	18.7	0.00	0.00	0.01	0.26	97.3	71.7
06JY30	Ol	40.5	0.00	0.05	0.04	9.72	0.15	48.9	0.12	0.00	0.00	0.32	99.8	90.1
	Opx	54.6	0.14	4.65	0.34	6.25	0.13	32.2	0.96	0.13	0.01	0.09	99.5	90.3
	Cpx	50.8	0.55	5.90	0.65	3.03	0.08	15.5	21.5	1.21	0.01	0.07	99.3	90.2
	Sp	0.03	0.30	53.9	10.0	11.8	0.12	20.7	0.00	0.01	0.01	0.30	97.2	75.9
06JY31	Ol	40.3	0.00	0.04	0.06	9.01	0.11	48.7	0.13	0.06	0.02	0.35	98.8	90.7
	Opx	55.3	0.10	3.38	0.60	5.86	0.12	32.7	0.94	0.11	0.00	0.11	99.2	90.9
	Cpx	52.4	0.30	4.51	1.09	2.75	0.09	15.9	21.6	1.10	0.01	0.07	99.8	91.2
	Sp	0.04	0.27	40.9	24.8	13.5	0.17	18.5	0.02	0.01	0.00	0.26	98.5	71.1
06JY34	Ol	41.0	0.00	0.02	0.05	8.11	0.13	49.6	0.07	0.01	0.00	0.34	99.3	91.7
	Opx	56.0	0.03	2.77	0.58	5.06	0.11	33.5	0.69	0.16	0.01	0.09	99.1	92.3
	Cpx	52.5	0.27	4.48	1.56	2.34	0.09	15.8	20.6	1.67	0.01	0.06	99.3	92.4
	Sp	0.01	0.27	37.3	29.2	12.0	0.17	18.2	0.03	0.04	0.01	0.18	97.4	73.1

Abbreviations: ol = olivine, opx = orthopyroxene, cpx = clinopyroxene, sp = spinel; Mg# = $100 \cdot \text{Mg}^{2+} / (\text{Mg}^{2+} + \text{Fe}^{2+})$.

data. For these standards our $\delta^7\text{Li} = -3.6 \pm 0.7$, $+10.6 \pm 0.9$, $+4.9 \pm 0.8$ and $+4.8 \pm 0.8$, respectively, consistent with the recommended values ($\delta^7\text{Li} = -4.1$, $+10.5$, $+4.4$ and $+4.2$, respectively) with analytical error. The external 2σ errors of the isotope compositions for both the standards and the samples are less than 2‰, with the majority less than 1.5‰ (Table 2). Li isotope ratios are given as $\delta^7\text{Li} (= [({}^7\text{Li}/{}^6\text{Li})_{\text{sample}} / ({}^7\text{Li}/{}^6\text{Li})_{\text{L-SVEC}} - 1] \times 1000)$ relative to the standard NIST L-SVEC with ${}^7\text{Li}/{}^6\text{Li}$ of 12.0192. Four to six minerals were measured for the Li and $\delta^7\text{Li}$ profiles for each sample. Note that the measured minerals are generally homogeneous in major elemental compositions at a mineral scale and show a limited range of olivine Mg# numbers (89.8–91.7; Table 1) similar to the standard ones. Thus matrix effects can be ignored in our measurement and the obtained data are reliable.

5. Results

5.1. Major elements

Olivine, opx, cpx and spinel in these xenoliths are generally homogeneous at a mineral scale by comparisons of core and rim analyses by EPMA. Thus the average composition is presented in Table 1 and plotted in Fig. 3. These xenoliths contain olivine with Fo = 89.8–91.7. Their opx have Mg# in the range of 90.3–92.3, Al₂O₃ of 2.77–4.65% and CaO of 0.69–0.96%. Clinopyroxenes have Mg# of 90.1–92.4, Al₂O₃ of 4.48–5.90% and Cr₂O₃ of 0.65–1.56%. The harzburgite (sample 06JY34) has the most magnesian olivine (Fo = 91.7) among these samples (Fig. 3), with compositions similar to the Fanshi harzburgite xenoliths and Hebi high-Mg# peridotite xenoliths (Fanshi and Hebi are two Cenozoic basaltic fields in the Central Zone of the NCC, Fig. 1). These were considered as the residues of ancient lithospheric mantle (Tang et al., 2008, 2011; Zheng et al., 2001). In contrast, the pyroxenite and lherzolite samples contain less magnesian olivine (Fo = 89.8–90.7), but high in Al₂O₃ contents of opx and cpx (Fig. 3), similar to the xenoliths of Hannuoba peridotites, Fanshi lherzolites, Hebi low-Mg# peridotites and those from many other localities in eastern China and other Phanerozoic volcanic areas of the world. All of the samples have mineral compositions similar to the peridotite xenoliths previously reported from Longgang (Fig. 3) (Xu et al., 2003).

5.2. Li abundances and isotopic compositions

For the pyroxenite (06JY30), the striking feature (Fig. 4) is high Li abundances in olivines (3.1–3.4 ppm) with variable $\delta^7\text{Li}$ ranging

from +1.5 to +5.8 (Table 2). Clinopyroxenes in 06JY30 show a style of $\delta^7\text{Li}$ zonation similar to that in 06JY06, with low $\delta^7\text{Li}$ in the rims and high $\delta^7\text{Li}$ in the cores.

In the cpx-poor lherzolite samples (06JY02 and 06JY06), olivines have a restricted range of Li (1.8–2.3 ppm) and $\delta^7\text{Li}$ (−1.9 to +0.8) (Table 2, Fig. 4). Opx grains show large variations in Li concentrations (0.8–1.9 ppm) and $\delta^7\text{Li}$ (−2.4 to +5.7). A profile analysis across a ~5 mm diameter opx in 06JY02 shows pronounced elemental and isotopic zonation, with a W-shaped $\delta^7\text{Li}$ zoning pattern (Fig. 5). In contrast, small opx grains show relatively weak zonation. Clinopyroxenes show very weak zonation of Li but clear zonation of $\delta^7\text{Li}$, variable from −0.8 to +3.6 from the rim to the core of cpx2 in 06JY06. In general, mineral $\delta^7\text{Li}$ in 06JY06 are low in the rims and high in the cores, except opx2 with $\delta^7\text{Li}$ varying from +3.9 at one side to −0.8 at another side (Fig. 5).

In the cpx-rich lherzolite (06JY31), olivines are enriched in Li (2.8–3.1 ppm) and variable in $\delta^7\text{Li}$ ranging from +2.2 to +5.4, very similar to those in pyroxenite (06JY30, Fig. 4). The $\delta^7\text{Li}$ of opx (+6.8 to +9.9) and cpx (+4.5 to +7.9) (Table 2) are higher than in coexisting olivine $\delta^7\text{Li}$ and higher in the rims and lower in the cores (Fig. 5).

In the harzburgite (06JY34), olivines are almost homogeneous (Fig. 5) in Li abundances but show zonation of low $\delta^7\text{Li}$ (−0.9 to −4.0). Opx and cpx are also homogeneous in Li contents (1.3–1.5 ppm in opx, 0.7–0.8 ppm in cpx) and variable in $\delta^7\text{Li}$ ranging from −0.2 to +5.7 in opx1, from −2.1 to +0.2 in opx2, and from +1.3 to +7.2 in cpx. One small cpx grain (cpx1) has Li abundance of 0.8 ppm and $\delta^7\text{Li}$ of +1.0 in the core (Table 2), similar to the rim compositions of coexisting cpx2 (larger than cpx1 in grain size).

In summary, Li abundances in olivines in each xenolith are generally homogeneous and higher than those of coexisting pyroxenes, with cpx having lower Li contents than coexisting opx (Fig. 4). However, $\delta^7\text{Li}$ in olivines are generally lower than those of coexisting pyroxenes. Among the samples, harzburgite is the lowest and pyroxenite is the highest in Li abundances and $\delta^7\text{Li}$ of olivines. There is a positive correlation between Li abundances and $\delta^7\text{Li}$ of olivines in these xenoliths (Fig. 6). Compared to the reported data for olivines from worldwide peridotites, olivines in the Longgang harzburgite and cpx-poor lherzolite xenoliths are very low in $\delta^7\text{Li}$. Pyroxenite and cpx-rich lherzolite xenoliths from the Longgang have intermediate Li abundances and $\delta^7\text{Li}$ of olivines, close to those of normal mantle (Fig. 6).

Table 2
In situ analytical data by SIMS for the xenoliths.

Sample	Mineral	Position	Li/ppm	$\delta^7\text{Li}/\%$	2σ		
06JY02	Olivine1	Rim	1.79	-1.67	0.82		
		Core	1.87	-1.90	0.91		
	Opx1	Rim	0.83	2.08	1.34		
			0.74	4.31	1.04		
		1.26	4.18	1.11			
		1.17	1.62	1.14			
		1.24	0.49	1.35			
		1.19	0.79	1.09			
		1.19	2.62	0.96			
		Core	1.27	2.98	1.07		
			1.20	2.82	1.00		
			1.23	3.23	0.76		
			1.21	-0.27	0.73		
			1.26	-0.19	0.97		
			1.16	3.23	0.98		
			0.78	5.37	1.30		
		Opx2	Rim	1.46	1.68	1.01	
	Rim		1.84	1.48	1.16		
	Opx3	Core	1.78	-0.41	1.10		
		Rim	1.14	5.65	1.20		
	Cpx1	Core	1.64	1.09	0.71		
		Rim	0.74	0.83	1.34		
	Cpx2	Core	0.66	0.76	1.24		
		Rim	1.00	0.51	1.53		
	06JY06	Olivine1	Core	1.01	-3.72	1.28	
			Rim	0.98	-0.83	1.06	
		Olivine2	Rim	2.19	-1.00	0.86	
				2.13	-0.72	0.52	
			2.27	0.79	0.56		
			1.57	-0.08	0.74		
			1.93	-1.34	0.58		
			2.02	-0.57	0.60		
			2.08	1.37	0.73		
2.18			-0.68	0.67			
Opx1			Rim	1.93	-2.39	0.93	
				1.83	2.36	0.84	
			1.82	0.50	0.81		
			Core	1.58	3.87	0.94	
				1.80	2.44	0.65	
				1.79	2.40	0.68	
				1.72	1.42	1.15	
1.70		2.55		0.79			
1.67		1.17	1.01				
1.64		-0.06	0.81				
Cpx1		Rim	1.68	-0.25	0.98		
			1.61	1.03	0.86		
		1.66	0.18	1.05			
		1.68	-0.81	0.86			
		1.66	0.11	0.41			
		1.18	2.57	1.22			
		1.18	4.02	0.91			
		Cpx2	Core	1.02	3.41	0.89	
			Rim	1.06	-0.80	1.16	
			Core	1.00	3.60	1.02	
		06JY30	Olivine1	Rim	3.28	1.47	0.63
				Core	3.27	2.34	0.74
			Olivine2	Core	3.41	1.88	0.77
	Rim			3.07	5.78	0.83	
	Core			3.29	2.64	0.68	
Cpx1	Rim			1.47	1.35	0.73	
	Core			1.31	6.00	0.97	
	Rim		1.23	7.81	1.09		
Cpx2	Core		0.59	-3.59	1.71		
	Rim		0.75	1.07	1.48		
06JY31	Olivine1		Core	1.04	5.11	1.50	
			Rim	1.15	5.40	1.22	
	Olivine2		Rim	2.76	2.56	0.51	
				2.90	5.41	0.85	
			2.89	2.19	0.64		
			2.95	3.73	0.51		
			3.03	3.77	0.50		
			3.12	3.54	0.52		
			2.76	3.98	0.66		
			Olivine3	Rim	2.76	3.98	0.66
					2.76	3.98	0.66
				2.76	3.98	0.66	
				2.76	3.98	0.66	
				2.76	3.98	0.66	
				2.76	3.98	0.66	
				2.76	3.98	0.66	

Table 2 (continued)

Sample	Mineral	Position	Li/ppm	$\delta^7\text{Li}/\%$	2σ	
06JY34	Olivine1	Core	3.07	4.58	0.79	
			2.78	2.24	0.66	
		Rim	2.85	3.64	0.67	
			1.17	9.89	0.59	
			1.32	7.01	0.58	
		Cpx1	Core	1.41	7.92	0.91
				1.46	6.80	0.92
			Rim	1.38	7.60	1.30
				1.04	4.50	1.08
				1.02	7.13	1.27
				0.97	6.95	1.16
				1.12	7.94	0.89
		Cpx2	Core	0.95	5.16	0.86
			Rim	0.90	6.77	0.79
		Olivine2	Rim	1.61	-3.14	0.81
	1.41			-4.03	0.95	
	Core		1.47	-2.97	0.93	
			1.51	-0.90	1.06	
			1.52	-3.22	0.62	
	Opx1		Rim	1.25	4.08	3.55
			Core	1.47	-0.22	1.21
			Core	1.49	5.74	1.57
	Opx2		Rim	1.30	-2.14	0.90
			Core	1.20	-0.59	1.10
			Core	1.30	0.23	1.16
	Cpx1		Core	0.77	1.00	1.12
			Rim	0.87	1.26	1.10
			Core	0.71	7.20	1.30
	Cpx2		Core	0.67	2.18	1.24
		Core	0.67	2.18	1.24	
		Core	0.67	2.18	1.24	

6. Discussion

6.1. Major element geochemistry and origin of mantle xenoliths

Extraction of mafic to ultramafic melts from a fertile peridotite with a primitive mantle composition will result in a residue depleted in basaltic components such as Al_2O_3 , CaO , and Na_2O . Therefore, typical Archean cratonic lithospheric mantle are highly refractory ($\text{Fo} > 92.5$) and depleted in basaltic components due to high-degree melting. In contrast, most Proterozoic and Phanerozoic lithospheric mantle are moderately depleted compared with primitive mantle (Beyer et al., 2006; O'Reilly et al., 2001). Thus, the variation in major-element compositions from fertile ($\text{Fo} \sim 90$) to refractory ($\text{Fo} \sim 92$) in the Longgang xenoliths (Fig. 3) may indicate that they formed by different degrees of melting possibly at different times. The harzburgite (06JY34) shows refractory mineral compositions ($\text{Fo} \sim 92$), similar to the Fanshi harzburgites and Hebi high-Mg# peridotites, which were considered as the residues of Archean lithospheric mantle (Tang et al., 2008, 2011; Zheng et al., 2001, 2006). Therefore, the Longgang harzburgite may represent the residues of ancient lithospheric mantle beneath this region. However, petrologic, trace element and Sr–Nd isotopic studies suggested that the Longgang harzburgites most likely resulted from melt–rock reaction rather than being simple residues of partial melting (Xu et al., 2003). This is consistent with published data for two refractory Longgang peridotites with Re-depletion ages of 0.6 and 0.4 Ga (Wu et al., 2003). This has been interpreted as evidence for melt–rock reaction, which may have resulted in lowering of the Re-depletion of the peridotites by introduction of Re (Beyer et al., 2006; Zhang et al., 2008a, 2009a).

Compared to the harzburgites, the Longgang pyroxenite and lherzolites are fertile in terms of their major elemental composition (low Fo, high Al_2O_3 ; Fig. 3) and are similar to those of the Hannuoba peridotites, Fanshi lherzolites, Hebi low-Mg# peridotites and the xenoliths from Cenozoic basalts of Shanwang (Zheng et al., 1998), Nūshan (Xu et al., 1998) and Qixia (Gao et al., 2002) in eastern China (Fig. 1). The latter were considered as the samples of newly accreted lithospheric mantle. Hence, the Longgang pyroxenite and

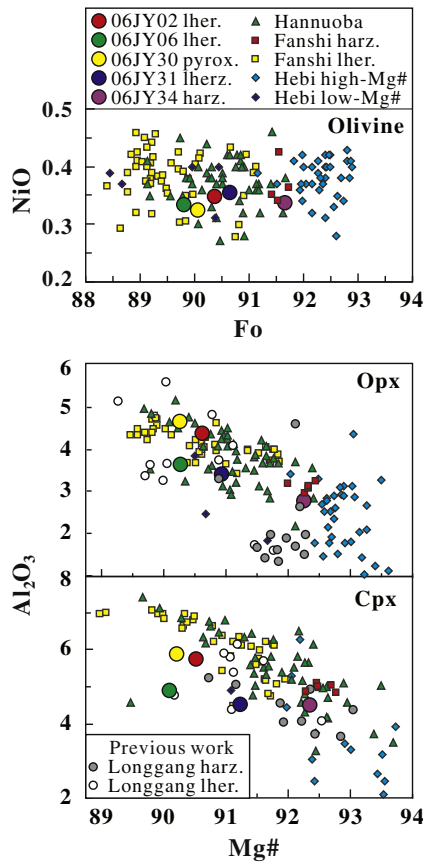


Fig. 3. Mineral compositions for the mantle xenoliths from the Longgang compared to the data of peridotite xenoliths from the Hannuoba (Fan and Hooper, 1989; Rudnick et al., 2004; Song and Frey, 1989; Tang et al., 2007, 2011; Tatsumoto et al., 1992; Zhang et al., 2009a), Fanshi (Tang et al., 2008, 2011; Xu et al., 2008b) and Hebi (Tang et al., 2011; Zheng et al., 2001, 2006) localities on the North China Craton. Previous work for Longgang harzburgite and lherzolite xenoliths are from Xu et al. (2003).

lherzolites may represent the newly formed lithospheric mantle or fragments of old lithospheric mantle modified by refertilization (see below).

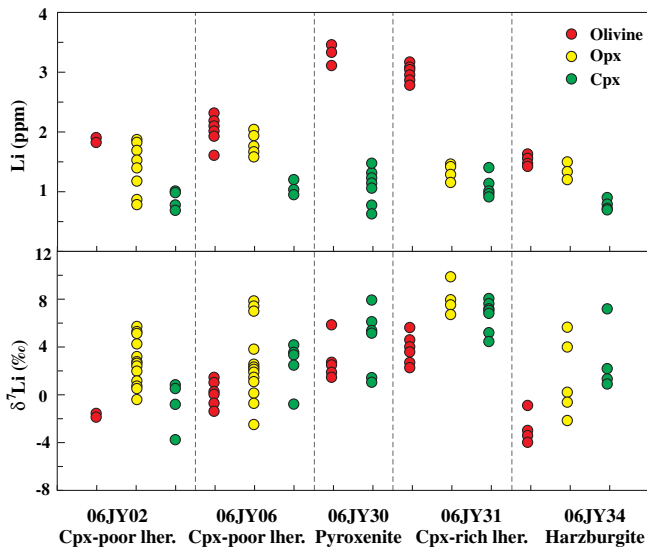


Fig. 4. Variations of Li abundances and $\delta^7\text{Li}$ of olivine, opx and cpx in the Longgang mantle xenoliths. The variation within the different minerals displays the different core and rim compositions of the minerals.

According to published data (Wu et al., 2003), four refractory peridotite xenoliths from the Longgang have Re-depletion ages of 1.0–1.2 Ga, suggesting a minimum Mesoproterozoic melt-depletion age of the lithospheric mantle beneath the northeastern NCC. The fact that most Longgang peridotites have Re-depletion ages older than 380 Ma (Wu et al., 2003) indicates that much of the lithospheric mantle formed before the Mesozoic and was not recently accreted. Thus, the mantle samples studied here likely formed during the Proterozoic even the Archean. The fertile compositions in the old peridotites may be the result of refertilization of lithospheric mantle because refertilization of an originally depleted source by melts rich in basaltic components can add Al, Ca and Fe to the peridotites, crystallizing cpx at the expense of olivine and opx (Griffin et al., 2009 and references therein). Many cases have been reported of transformation of highly refractory lithospheric mantle via refertilization, producing relatively fertile lherzolites and pyroxenites (e.g., Beyer et al., 2006; Carlson et al., 2004; Griffin et al., 2005; Stiefenhofer et al., 1997). Therefore, fertile lherzolite and pyroxenite xenoliths from the Longgang may be produced by refertilization via refractory peridotite–melt reaction. Olivine Fo (91.7) in the harzburgite is less than that (Fo > 92.5) of typical Archean cratonic mantle, indicating that the harzburgite is not a simple residue of partial melting but a product of melt–rock reaction, similar to those from the Archean cratons of Kaapvaal and Siberia (Kelemen et al., 1998).

6.2. Slab-derived Li isotopic signature and multistage metasomatism

Li, a moderately incompatible element (Brenan et al., 1998a, 1998b), has relatively high diffusivity, with ^6Li diffusion faster than ^7Li in mantle minerals (Coogan et al., 2005; Dohmen et al., 2010; Richter et al., 2003). Therefore, diffusion-driven kinetic fractionation has been invoked to account for Li isotopic variations in peridotites. Diffusive fractionation of Li isotopes can occur during peridotite–melt/fluid reaction before or coincident with the entrainment into host magmas and the transport of the mantle xenolith to the surface (e.g., Aulbach and Rudnick, 2009; Aulbach et al., 2008; Gallagher and Elliott, 2009; Halama et al., 2009; Ionov and Seitz, 2008; Kaliwoda et al., 2008; Pogge von Strandmann et al., 2011; Rudnick and Ionov, 2007). Thus, it is necessary to discuss the influence of late processes on Li abundances and isotopic compositions of minerals in mantle xenoliths before probing into the signature of lithospheric mantle represented by the xenolith.

6.2.1. Slab-derived Li isotopic signature in the harzburgite

It is known that diffusive fractionation of Li depends on both the type of minerals and their grain size. Small minerals are possibly better equilibrated within Li compared to large grains (e.g., Halama et al., 2009; Ionov and Seitz, 2008). Coarse-grained olivines (2–3 mm in diameter) in the harzburgite (06JY34) are generally homogeneous in Li abundances and show a restricted variation in $\delta^7\text{Li}$ (Fig. 5), suggesting no significant effect of diffusive fractionation on the compositions of olivines during their entrainment into the host magmas and transportation of the mantle xenolith to the surface. Moreover, low Li abundances and $\delta^7\text{Li}$ in olivines in 06JY34 further excludes a significant influence of the host magma because the host magma, derived from the asthenosphere, will elevate Li contents and $\delta^7\text{Li}$ of xenolithic olivines making them more similar to normal mantle. In addition, the Longgang harzburgite were entrained by pyroclastic rocks or scoriaeous lavas, which cooled almost instantaneously on eruption. The near-instant cooling of the xenolith “freezes” its Li inventories, preventing significant fractionation of Li isotopes due to diffusive re-distribution of Li during long cooling driven by temperature-controlled inter-mineral partition coefficients (Gallagher and Elliott, 2009; Gao et al., 2011; Ionov and Seitz, 2008; Kaliwoda et al., 2008). Therefore, low $\delta^7\text{Li}$ in olivines from the harzburgite reflect low $\delta^7\text{Li}$ of the lithospheric mantle represented by the xenolith. Similarly,

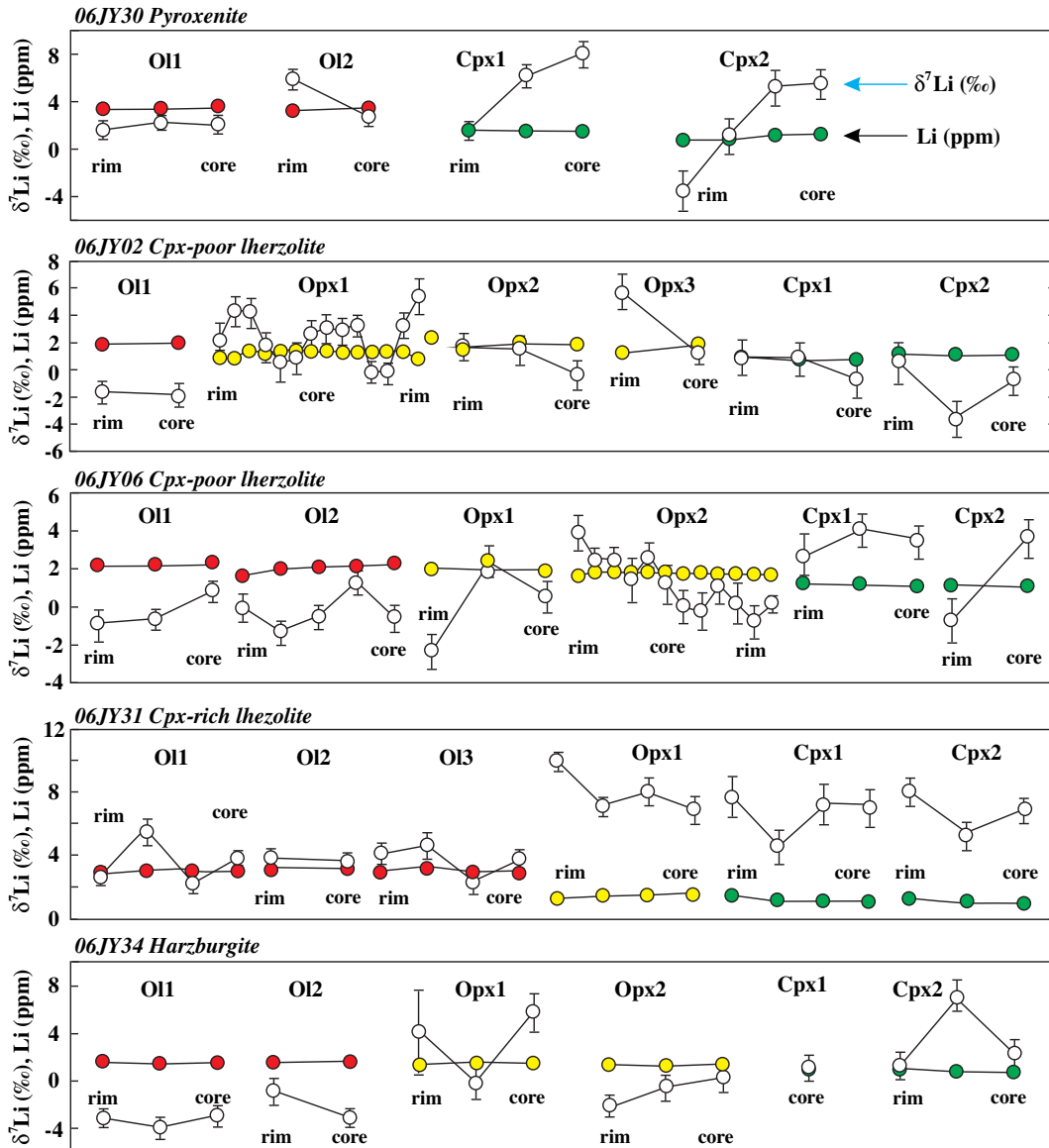


Fig. 5. Profiles of Li abundances and isotopic compositions of olivine, opx and cpx in the Longgang xenoliths analyzed by SIMS.

low $\delta^7\text{Li}$ in olivines in the lherzolites (Fig. 5) also imply the existence of low- $\delta^7\text{Li}$ lithospheric mantle beneath the Longgang region.

How did the low- $\delta^7\text{Li}$ lithospheric mantle form? There are two possible models (1) diffusion-driven kinetic fractionation of Li isotopes (e.g. Gallagher and Elliott, 2009; Gao et al., 2011; Ionov and Seitz, 2008; Jeffcoate et al., 2007; Kaliwoda et al., 2008; Lundstrom et al., 2005; Marschall et al., 2007; Rudnick and Ionov, 2007; Tang et al., 2011; Teng et al., 2006), (2) metasomatism of lithospheric mantle by low- $\delta^7\text{Li}$ melt/fluid (e.g., Agostini et al., 2008; Nishio et al., 2004; Tang et al., 2007).

If low $\delta^7\text{Li}$ in mantle olivines were produced by diffusion-driven fractionation of Li isotopes (possibility #1), olivines and pyroxenes must be enriched in Li. However, we observe that Li abundances in the minerals (1.5–1.6 ppm in olivine, 1.3–1.5 ppm in opx, 0.7–0.9 ppm in cpx; Table 2) of the harzburgite are close to, or slightly lower than, those of normal mantle (1–1.8 ppm in olivine, 0.5–1.3 ppm in pyroxene; Seitz and Woodland, 2000; Woodland et al., 2004). Furthermore, diffusion-driven kinetic fractionation of Li will produce lower $\delta^7\text{Li}$ in cpx than in coexisting olivines (Aulbach and Rudnick, 2009; Ionov and Seitz, 2008; Rudnick and Ionov, 2007; Tang et al., 2007) because the diffusivity of Li is much higher in

pyroxene than in olivine (Dohmen et al., 2010; Parkinson et al., 2007) and Li^6 diffuses faster than Li^7 . Since $\delta^7\text{Li}$ of the pyroxenes in the harzburgite is higher than those of coexisting olivines (Fig. 4) the first model cannot explain the Li abundances and isotopic compositions. Therefore the second model best explains the low $\delta^7\text{Li}$ in olivines, a model involving metasomatism of the harzburgite by low- $\delta^7\text{Li}$ melt/fluid.

Where does the low- $\delta^7\text{Li}$ melt/fluid originate? Based on experimental results, Wunder et al. (2006, 2007) suggest that subduction and concomitant dehydration of altered oceanic crust, containing chlorite and cpx with Li in six-fold coordination, release fluids enriched in ^7Li into the fore-arc mantle (Chan et al., 2002a; Elliott et al., 2003) and consequently introduce a low- $\delta^7\text{Li}$ component into the deeper mantle. This suggestion supports the model that low $\delta^7\text{Li}$ in eclogites were derived from Li isotopic fractionation during progressive metamorphism of the subducting slab (Zack et al., 2003). However, Marschall et al. (2007) estimate that purely solid–fluid processes will reduce $\delta^7\text{Li}$ in deeply subducted slabs by no more than 3‰. Based on the study of coupled Li and B isotope variations in western Anatolia Cenozoic lavas, Agostini et al. (2008) present an alternative interpretation of Li isotopes similar to the low $\delta^7\text{Li}$ eclogites of

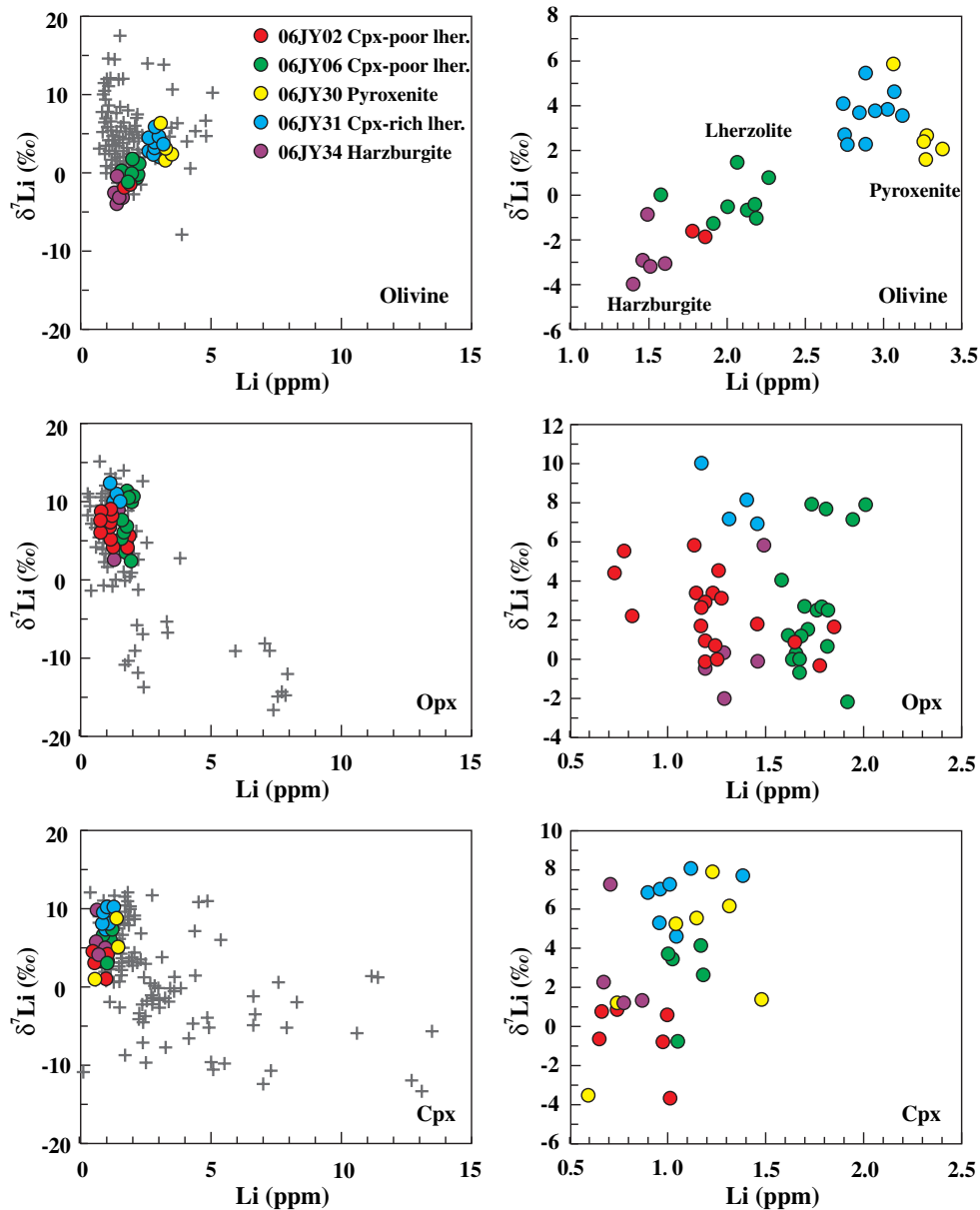


Fig. 6. Variation of Li abundances with $\delta^7\text{Li}$ of olivine, opx and cpx in the Longgang xenoliths compared to worldwide data.

Data sources: Nishio et al. (2004), Seitz et al. (2004), Magna et al. (2006), Jeffcoate et al. (2007), Rudnick and Ionov (2007), Halama et al. (2009), Tang et al. (2007, 2011), Wagner and Deloule (2007), Ionov and Seitz (2008), Aulbach and Rudnick (2009), Zhang et al. (2010).

Marschall et al. (2007) and argue that Li isotopic fractionation during fluid releasing will not result in extensive removal of Li from the slab due to its moderate incompatibility (Brenan et al., 1998a, 1998b). The ongoing solid–fluid exchanges could lower the $\delta^7\text{Li}$ of the slab extensively at shallow depths where the slab is being comprehensively devolatilized. Thus the very dehydrated slab will have a low $\delta^7\text{Li}$, which will be preserved in the local mantle (Agostini et al., 2008). Considering the low $\delta^7\text{Li}$ in eclogites (Halama et al., 2011; Marschall et al., 2007; Zack et al., 2003) and variable $\delta^7\text{Li}$ in the upper continental crust (Teng et al., 2004), upper oceanic crust and sediments (e.g., Bouman et al., 2004; Chan et al., 2002b; Moriguti and Nakamura, 1998), we suggest that the eclogites could be a possible source for the low- $\delta^7\text{Li}$ melt.

The main reason we favor the second model is that the low- $\delta^7\text{Li}$ metasomatic melt/fluid was likely derived from recycled dehydrated Pacific plate. Since the Late Mesozoic, eastern China has become an important part of the circum-Pacific tectonic–magmatic zone (Wu et al., 2005). The westward subduction of the Pacific plate could have generated a large mantle wedge above the subducted oceanic

slab and significantly affected the lithospheric mantle beneath eastern China (Zhao et al., 2007). Geochemical (Li and Li, 2007; Liu et al., 2010; Sun et al., 2007; Zhang et al., 2008b, 2009b) and geophysical evidence (Zhao et al., 2007; Zhu and Zheng, 2009) suggest an important role for Pacific plate subduction in the lithospheric evolution of eastern China. For example, Cenozoic basalts from the eastern NCC provide a petrological record of modification of the lithospheric mantle by high temperature sub-lithospheric melts (Zhang et al., 2009b). Longgang is located in the northeastern China (Fig. 1) and thus the lithospheric mantle beneath this region could have been significantly affected by subduction of the Pacific plate since the Late Mesozoic. In addition, the Longgang harzburgite was likely the product of melt–rock reaction as discussed above and it could be produced by a two step process, (1) high Mg#, low opx peridotite was created by high-degree melting, and (2) the depleted residue was enriched in opx by reaction with SiO_2 -rich melts derived mainly by melting of eclogitic basalt and sediment in a subduction zone (Kelemen et al., 1998). As a result, we consider that low $\delta^7\text{Li}$ in the lithospheric mantle

could be generated by reaction between peridotites and low- $\delta^7\text{Li}$ melt derived from recycled dehydrated Pacific plate since the Late Mesozoic.

Now the question is whether low $\delta^7\text{Li}$ in the mantle could survive diffusion over about 100 Ma (since the Late Mesozoic). Many previous studies (e.g., Aulbach and Rudnick, 2009; Gallagher and Elliott, 2009; Halama et al., 2008; Ionov and Seitz, 2008; Jeffcoate et al., 2007; Parkinson et al., 2007; Rudnick and Ionov, 2007) suggested that Li isotopic heterogeneity in the mantle can be rehomogenized by diffusion over geologically short time periods due to fast diffusion of Li. Recently two mechanisms of Li diffusion have been proposed based on experimental observations (1) a fast (interstitial site) mechanism and (2) a slow (vacancy, octahedral site) mechanism (Dohmen et al., 2010). The fast mechanism of Li diffusion is unlikely to be dominant in most natural systems (oxidizing environments may be exception). The diffusion of Li when dominated by the slower mechanism is about an order of magnitude faster than diffusion of Fe, Mg and most other divalent cations in olivine. Such diffusion of Li in olivine will be much slower than the diffusion in cpx and plagioclase crystals at the same conditions (Dohmen et al., 2010). This observation could reflect a temperature dependency within the diffusion of Li in olivine versus cpx (Gao et al., 2011; Kaliwoda et al., 2008; Wunder et al., 2006). In situ investigation of a natural olivine xenocryst showed that the effective diffusion coefficients of many elements (e.g., Li, Ca and Y) in olivine fell within a factor of three of the Mg–Fe interdiffusion coefficient (Qian et al., 2010). This observation is in agreement with the experimental results that the diffusion rate of Li into olivine is similar to that of the divalent cations (Dohmen et al., 2010; Parkinson et al., 2007; Spandler and O'Neill, 2010). Moreover, a kilometer-thick altered oceanic crust loses most of its Li excess in a matter of a few tens of millions years, consistent with Halama et al. (2008). However it could preserve $^7\text{Li}/^6\text{Li}$ that was distinctively higher than the ambient mantle over a time period in excess of 1.5 Ga (Vlastélic et al., 2009) even though the calculation is based on the fast mechanism of diffusion. In addition, large-scale heterogeneity could survive diffusion of Li over time scales longer than small-scale heterogeneity (Vlastélic et al., 2009). Therefore, low $\delta^7\text{Li}$ of the lithospheric mantle caused by low- $\delta^7\text{Li}$ melt derived from the subducted Pacific plate could survive diffusion over about 100 Ma and thus can be recorded by the mantle xenoliths.

In theory, if a peridotite with $\delta^7\text{Li}$ of normal mantle is modified and in equilibrium with low- $\delta^7\text{Li}$ melt, its minerals (including olivine, opx and cpx) should be lower in $\delta^7\text{Li}$ and higher in Li contents than those of normal mantle. However, the pyroxenes in 06JY34 are higher in $\delta^7\text{Li}$ than coexisting olivines and the component minerals have Li abundances similar to the normal mantle (Fig. 4). These observations suggest a high-degree partial melting after metasomatism of the peridotite by low- $\delta^7\text{Li}$ melt, followed by recent reaction between peridotite and asthenosphere-derived melt. The peridotite could become fertile and enriched in Li due to reaction with low- $\delta^7\text{Li}$ melt, and then the Li concentrations of minerals in the peridotite were lowered by high-degree melting such that they approached those of normal mantle. In addition high-degree melting could convert the fertile peridotite into relatively refractory harzburgite. This is consistent with the Re–Os isotopic data of Longgang peridotite xenoliths, suggesting extensive melt extraction of some portions of the lithospheric mantle during the Phanerozoic (Wu et al., 2003). Pyroxenes in the harzburgite have variable $\delta^7\text{Li}$ within the range of normal mantle, indicative of recent reaction between the peridotite and asthenosphere-derived melt. This is in accordance with the petrologic and geochemical signatures of Longgang harzburgites, reflecting interaction between the lithosphere and asthenosphere during lithospheric thinning (Xu et al., 2003).

6.2.2. Varying $\delta^7\text{Li}$ in the lherzolites and pyroxenite

The cpx-poor lherzolites 06JY02 and 06JY06 are similar in Li abundances and isotopic compositions in minerals (Fig. 5). They

successively experienced low- $\delta^7\text{Li}$ melt metasomatism, partial melting and recent metasomatism by asthenospheric melt, similar to the harzburgite. Olivines in the lherzolites are low in $\delta^7\text{Li}$ and have Li abundances close to that of normal mantle, implying partial melting of the lherzolites after low- $\delta^7\text{Li}$ melt metasomatism. Likewise, Li abundances and $\delta^7\text{Li}$ in opx and cpx from the lherzolites suggest recent asthenospheric melt–peridotite reaction after low- $\delta^7\text{Li}$ melt metasomatism and partial melting. High Li concentrations and low $\delta^7\text{Li}$ in the rim of cpx in lherzolite 06JY06 reflect diffusive influx of Li during recent peridotite–melt interaction and the entrainment by host magma. The W-shape $\delta^7\text{Li}$ zoning pattern in coarse-grain opx in 06JY02 (Fig. 5) records multistage metasomatism. The flat core of the opx is similar to normal mantle in Li abundance and $\delta^7\text{Li}$, indicating that the opx core records primary compositions of normal mantle. Rims of the opx are higher in $\delta^7\text{Li}$ and lower in Li abundance than the core, and have Li abundances and $\delta^7\text{Li}$ close to those of normal mantle, possibly reflecting recent metasomatism by asthenospheric melt. The transitional zones between both the rims and the core of the opx have $\delta^7\text{Li}$ lower than those of normal mantle, but their Li abundances are close to that of normal mantle. These observations suggest partial melting after an early metasomatism of the lherzolite by low- $\delta^7\text{Li}$ melt.

The cpx-rich lherzolite (06JY31) and pyroxenite (06JY30) are similar in Li contents and $\delta^7\text{Li}$ in olivine, opx and cpx, respectively. Li abundances in olivines are clearly higher than those of normal mantle olivine, while $\delta^7\text{Li}$ in olivines are in the range of normal mantle, reflecting metasomatism of the peridotites by asthenospheric melt.

6.2.3. Comparison between published data and our results

Compared to published data for worldwide peridotites (Fig. 6), the minerals in the Longgang xenoliths show limited variations in Li abundances and $\delta^7\text{Li}$. A striking feature is that $\delta^7\text{Li}$ in olivines from the Longgang harzburgite and cpx-poor lherzolites are lower than those published for olivines from worldwide peridotites. Olivines in cpx-rich lherzolite and pyroxenite have normal mantle-like $\delta^7\text{Li}$, close to most of the published data. These observations suggest that the lithospheric mantle beneath the Longgang region is peculiar due to metasomatism by low- $\delta^7\text{Li}$ melt derived from the recycled oceanic plate. A generally positive correlation between $\delta^7\text{Li}$ and Li abundances in the Longgang olivines (Fig. 6) further indicates mixing between low- $\delta^7\text{Li}$ melt and normal mantle. In contrast, the scatter on the plots of $\delta^7\text{Li}$ and Li contents in opx and cpx reflects the susceptibility of pyroxenes to recent diffusive fractionation of Li isotopes during the entrainment and transportation of the xenoliths by host magma due to faster diffusion of Li in pyroxene than that in olivine.

6.3. Constraints on the transformation of lithospheric mantle beneath the NCC

The mechanism responsible for lithospheric thinning and compositional change is a long-standing puzzle. Up to date, there are at least four models invoked to account for the transformation of lithospheric mantle beneath the NCC, including: (1) delamination of lithospheric mantle and even lower crust (Deng et al., 2004; Gao et al., 2004; Wu et al., 2003), (2) thermo-mechanical and chemical erosion (Menzies and Xu, 1998; Xu, 2001), (3) partial replacement of old refractory lithospheric mantle by newly accreted fertile mantle (Zheng et al., 2001, 2006), and (4) transformation by peridotite–melt reaction (Zhang, 2005; Zhang et al., 2002, 2007, 2009a).

Whilst these models can solve different aspects of lithospheric thinning and compositional changes related to the destruction of the NCC, some of the evidence is equivocal. For example, in models of delamination, the lithospheric mantle should be accreted in the Mesozoic or Cenozoic due to the complete removal of Archean/Precambrian lithosphere by delamination (Gao et al., 2004; Wu et al., 2003), but this is in conflict with the presence of numerous

peridotite xenoliths from the eastern NCC with Proterozoic Re depletion ages and/or Os model ages (Chu et al., 2009; Gao et al., 2002; Wu et al., 2003, 2006; Xu et al., 2008a, 2008b; Zhang et al., 2009a). Therefore, the delamination model cannot explain the coexistence of old and young lithospheric mantle.

Due to partial removal of old lithospheric mantle in the model of thermo-mechanical and chemical erosion (Xu, 2001) and replacement (Zheng et al., 2001), it seems certain that the two models can properly explain the coexistence of old and young lithospheric mantle (Griffin et al., 1998; Menzies and Xu, 1998). However, the residue of old lithospheric mantle should have Archean ages (Gao et al., 2002; Wu et al., 2006; Zhang et al., 2008a) and highly refractory compositions typical of Archean cratonic mantle as that existed prior to its thinning. This is also inconsistent with the observations that the present lithospheric mantle beneath the NCC has a complex age structure, with dominant Proterozoic ages, but does not appear to retain rocks with Archean melt depletion ages (Chu et al., 2009; Wu et al., 2006; Zhang et al., 2009a). In addition the peridotites believed to be relict refractory mantle that are sampled by the Cenozoic volcanic rocks have a dominant $Fo=92$, less than those of typical Archean cratonic mantle ($Fo>92.5$). Hence, the two models have limitations.

In contrast, the model of transformation of lithospheric mantle by peridotite–melt reaction may be a more appropriate model because refertilization by peridotite–melt reaction can explain the compositional change of the lithospheric mantle and the juxtaposition of young and old lithospheric mantle (Zhang et al., 2008a, 2009a). The characteristics of Li abundances and δ^7Li in olivine, opx and cpx from the Longgang xenoliths reflect multistage metasomatism of the lithospheric mantle; the early-stage metasomatic melt/fluid could be derived from subducted oceanic plate and later-stage metasomatic melt was derived from the asthenosphere. These processes triggered petrological, elemental and isotopic changes in the lithospheric mantle beneath the northeastern NCC. The character of the lithospheric mantle also changed from Archean cratonic type into oceanic type. Based on these results and previous observations of geochemistry and geophysics (Li and Li, 2007; Wu et al., 2005; Zhao et al., 2007; Zhu and Zheng, 2009; Zhu et al., 2011), reactions between peridotites and melts/fluids derived from subducted Pacific plate may be the essential for transformation of the lithospheric mantle and the destruction of the eastern NCC.

7. Conclusions

Mantle xenoliths from the Longgang, northeastern NCC include spinel-facies cpx-poor and cpx-rich lherzolites, harzburgite and pyroxenite. Lherzolite and pyroxenite xenoliths are fertile in major elemental compositions relative to the harzburgite. Olivines in the harzburgite and cpx-poor lherzolites have generally homogeneous Li contents, in the range of normal mantle, and low δ^7Li (–0.9 to –4.0 in harzburgite, –1.9 to +0.8 in cpx-poor lherzolites), while those in cpx-rich lherzolite and pyroxenite have high δ^7Li (+2.2 to +5.4 in cpx-rich lherzolite, +1.5 to +5.8 in pyroxenite) and Li contents (2.8–3.4 ppm). Pyroxenes generally have lower Li contents and higher δ^7Li than coexisting olivines. These characteristics and in situ analyzed profiles of Li contents and δ^7Li in minerals from the xenoliths, coupled with published geochemical and geophysical observations, allow us to draw the following conclusions:

- (1) Low- δ^7Li signature, particularly evident in the harzburgite, reflects the presence of a slab-derived Li isotopic signature in the lithospheric mantle beneath the northeastern NCC, which must have experienced multistage metasomatism. The older metasomatic melt was low in δ^7Li and probably derived from recycled oceanic plate whereas the younger metasomatic melt may have been derived from the asthenosphere.

- (2) The lithospheric mantle beneath the NCC has been significantly refertilized by peridotite–melt reactions. The subducted Pacific plate-derived melts/fluids could play a crucial role in the destruction of the NCC.

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