

## The genesis of mantle-derived sapphirine

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### ABSTRACT

Sapphirine, a typical ultrahigh-temperature metamorphic mineral, is rarely found in mantle xenoliths. Here we report the occurrence and characteristics of sapphirine in a mantle-derived xenolith from the Cenozoic basalts of Hannuoba in the North China Craton. The xenolith consists of a clinopyroxene, spinel, and sapphirine assemblage, with the sapphirine occurring as the reaction rim surrounding spinel. The mineral compositions of this sample are all characterized by high Mg contents, similar to those of minerals from other sapphirine-bearing rocks reported from high-Mg-Al granulites elsewhere in the world. Clinopyroxene is relatively rich in Al and Ca in comparison to pyroxene in peridotite and pyroxenite xenoliths in the Hannuoba basalts, as well as in global mafic and felsic granulites in other terranes, a feature that is consistent with the bulk composition. The *P-T* compilations from both experimental and natural rock data show a restricted stability field for the coexisting clinopyroxene + spinel + sapphirine assemblage of around 8–15 kbar and 800–900 °C. The rare occurrence of sapphirine in a mantle-derived xenolith therefore suggests specific bulk composition, restricted *P-T* range and possible melt-peridotite interaction. Such conditions are best satisfied in a tectonic setting with basaltic magma underplating and interaction between the infiltrating melts and the wall-rock peridotite.

**Keywords:** Sapphirine, pyroxenite, *P-T* condition, North China Craton

### INTRODUCTION

Experimental studies and *P-T* estimates from natural assemblages and mineral phase equilibria modeling have constrained the stability of sapphirine under ultrahigh-temperature (UHT) and medium- to high-pressure conditions of 900–1100 °C and 10–18 kbar, corresponding to lower crustal and upper mantle depths (e.g., Ackermann et al. 1975; Sills et al. 1983; Powell and Holland 1985; Griffin and O'Reilly 1986; Christy 1989; Das et al. 2006; Sato et al. 2006; Brigida et al. 2007; Podlesskii et al. 2008; Kelsey 2008). Sapphirine has been reported in Mg-Al granulites subjected to extreme crustal metamorphism under UHT conditions in various regions of the world (e.g., Lal et al. 1987; Okay 1994; Santosh et al. 2007, 2009; Galli et al. 2011; Jiao and Guo 2011). The stability of sapphirine-bearing assemblages in metamorphic orogens exhumed from depth has been a topic of wide interest in characterizing metamorphic *P-T* conditions, tectonic architecture, and crustal evolution (e.g., Sills et al. 1983; Christy 1989; Okay 1994; Mattielli et al. 1996; Gregoire et al. 2001; Harley 2004; Santosh et al. 2007, 2009, 2012; Tsunogae and Santosh 2011; Tsunogae et al. 2011; Xiang et al. 2012).

Mantle-derived sapphirine has been rarely reported, such as in pyroxenite xenoliths from Stockdale, Kansas (Meyer and Brookins 1976), from Delegate, New South Wales, Australia (Griffin and O'Reilly 1986), the Beni Bousera massif, Morocco (Kornprobst et al. 1990), and the Ronda massif, Spain (Morishita et al. 2001). Sapphirine in these pyroxenites exhibits diverse

textural relationships despite the relatively consistent mineral assemblages, suggesting that sapphirine occurs in a wide range of *P-T* conditions and bulk compositions. Griffin and O'Reilly (1986) suggested that sapphirine is stable under upper-mantle conditions in Ca-Al-Mg-rich bulk compositions, whereas Okay (1994) suggested that the vast majority of sapphirine-bearing rocks are rich in Mg and Al and poor in Ca, consistent with the composition of the mineral. Sapphirine occurs as relics of a low-*P* mafic granulite assemblage preserved in garnet in Okay's (1994) study. Christy (1989) suggested that sapphirine might be a widespread constituent of basic rocks with high Mg/(Mg+Fe) ratio. The models proposed to explain the origin of sapphirine in such rocks include recycling of crustal material (Morishita et al. 2001) and exsolution of clinopyroxene (Kornprobst et al. 1990; Mattielli et al. 1996).

The Cenozoic Hannuoba basalts occur along the northern part of the Trans-North China Orogen, a Paleoproterozoic suture that amalgamates the Western and Eastern Blocks in the North China Craton. Abundant lower crustal and upper mantle xenoliths are found in the basalts and have been studied by various workers (Fan et al. 2001; Liu et al. 2010; Rudnick et al. 2004; Zhang 2009; Zheng et al. 2009; Zhang et al. 2011, 2012). However, the present study is the first report of sapphirine-bearing rocks from Hannuoba. We present detailed petrologic and geochemical data on the sapphirine-bearing clinopyroxenite from Hannuoba and compare our results with other occurrences in an attempt to characterize the origin of mantle-derived sapphirine, as well as suggest possible chemical discriminants between sapphirines of crustal and mantle origins.

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## GEOLOGICAL BACKGROUND AND SAMPLE DESCRIPTIONS

The Tertiary Hannuoba basalt field is located on the northern margin of the North China Craton (Fig. 1a) and has been dated at 22 to 14 Ma (Liu et al. 1990). It is best known for its abundance of diverse mantle and crustal xenoliths. The mantle-derived xenoliths vary in composition from peridotite (dunite, harzburgite, lherzolite, and wehrlite) to garnet- and spinel-facies pyroxenite (e.g., Chen et al. 2001; Xu 2002; Zhou et al. 2002; Rudnick et al. 2004; Liu et al. 2005; Tang et al. 2007; Zhang et al. 2012), whereas the crustal xenoliths are basically divided into mafic and felsic types (e.g., Fan et al. 2005; Liu et al. 2005; Zheng et al. 2009; Jiang and Guo 2010). These xenoliths are mainly found in three localities: Damaping, Zhouba, and Jieshaba (Fig. 1b).

The sapphire-bearing clinopyroxenite (JSB10-57) studied here was collected from Jieshaba. The gray-green xenolith within host basalt shows rounded edges and measures approximately  $25 \times 15 \times 5$  cm and has a medium-grained equigranular texture. There are only three kinds of rock-forming minerals in this

sample, namely clinopyroxene (~90% by mode), spinel (9%), and sapphire (1%). The clinopyroxene is sub-rounded and 1–3 mm in length (Fig. 2a). Some grains show well-developed cleavage along which orthopyroxene exsolution lamellae occur in some grains (Figs. 2b, 2c, and 2d). Some grain boundaries and fractures show a locally spongy-texture enclosing tiny ilmenite and glass (Figs. 2c and 2e). Spinel is optically light green, sub-hedral to anhedral in form, and occurs either as inclusions in, or as discrete grains between adjacent clinopyroxenes (Figs. 2b and 2d). Two types of spinel are identified: a fine-grained (<0.3 mm) variety and a second type (>0.2 mm) rimmed by sapphire (Figs. 2b–2f). Sapphire occurs as thin, pale-blue rims (<150  $\mu\text{m}$ ) grown around spinel grains. Sapphire is also found as symplectitic intergrowths with secondary spinel or clinopyroxene; occasionally inclusions of clinopyroxene are trapped within the sapphire (Fig. 2e), with a rim of secondary spinel surrounding the enclosed sapphire (Fig. 2f).

## BULK-ROCK CHEMISTRY AND MINERAL COMPOSITION

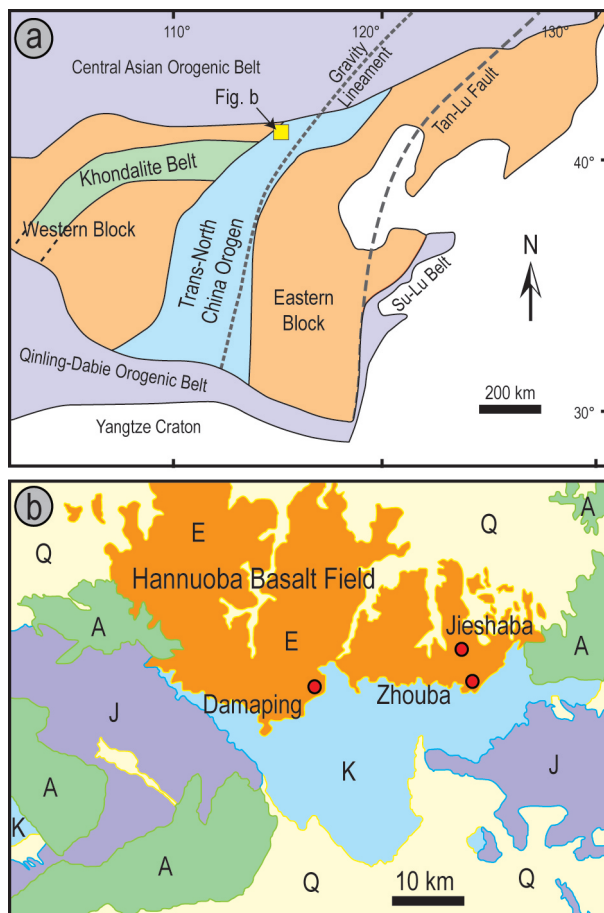
The bulk major element chemistry was obtained by X-ray fluorescence spectrometry (XRF) on fused glass disks using a Shimadzu XRF-1500 instrument. FeO content was analyzed by a chemical titration method and  $\text{Fe}_2\text{O}_3$  was then recalculated using total Fe and FeO. Mineral chemistries were measured by wavelength-dispersive spectrometry (WDS) using a JEOL JXA8100 electron probe operating at an accelerating voltage of 15 kV with 12 nA beam current, 5  $\mu\text{m}$  beam spot, and 10–30 s counting time. Both measurements were carried out at the Institute of Geology and Geophysics, Chinese Academy of Sciences.

The sapphire-bearing spinel clinopyroxenite is poor in  $\text{SiO}_2$  (40.3 wt%) and rich in  $\text{Al}_2\text{O}_3$  (20.4 wt%). It has a remarkably high Mg no. [ $\text{Mg no.} = 100 \times \text{Mg}/(\text{Mg} + \text{Fe}_{\text{total}})$ ] of 85.3 with  $\text{FeO}/\text{Fe}_2\text{O}_3$  ratio <1. The CaO content is 16.2 wt% (Table 1).

The clinopyroxenes in this sample are aluminous diopside ( $\text{Al}_2\text{O}_3$ : 5.74–10.4 wt%) with low  $\text{Cr}_2\text{O}_3$  contents (0.12–0.16 wt%). The clinopyroxene in the spongy rim shows higher  $\text{SiO}_2$ , MgO, CaO, and apparently lower  $\text{Al}_2\text{O}_3$  and  $\text{Na}_2\text{O}$  contents than those of the cores (Table 2). The clinopyroxene grain enclosed in sapphire shows lower FeO content (2.80 wt%) and higher Mg no. (90.0) relative to the discrete clinopyroxene grains. Clinopyroxene has variable  $\text{Fe}^{2+}/\text{Fe}^{3+}$  ratio, and one grain analyzed shows a low  $\text{Fe}^{2+}/\text{Fe}^{3+}$  ratio of <1 (Table 2). The orthopyroxene lamellae in clinopyroxene are rich in  $\text{Al}_2\text{O}_3$  (7.29 wt%) but poor in  $\text{SiO}_2$  (52.4 wt%) and low in Mg no. (87.7).

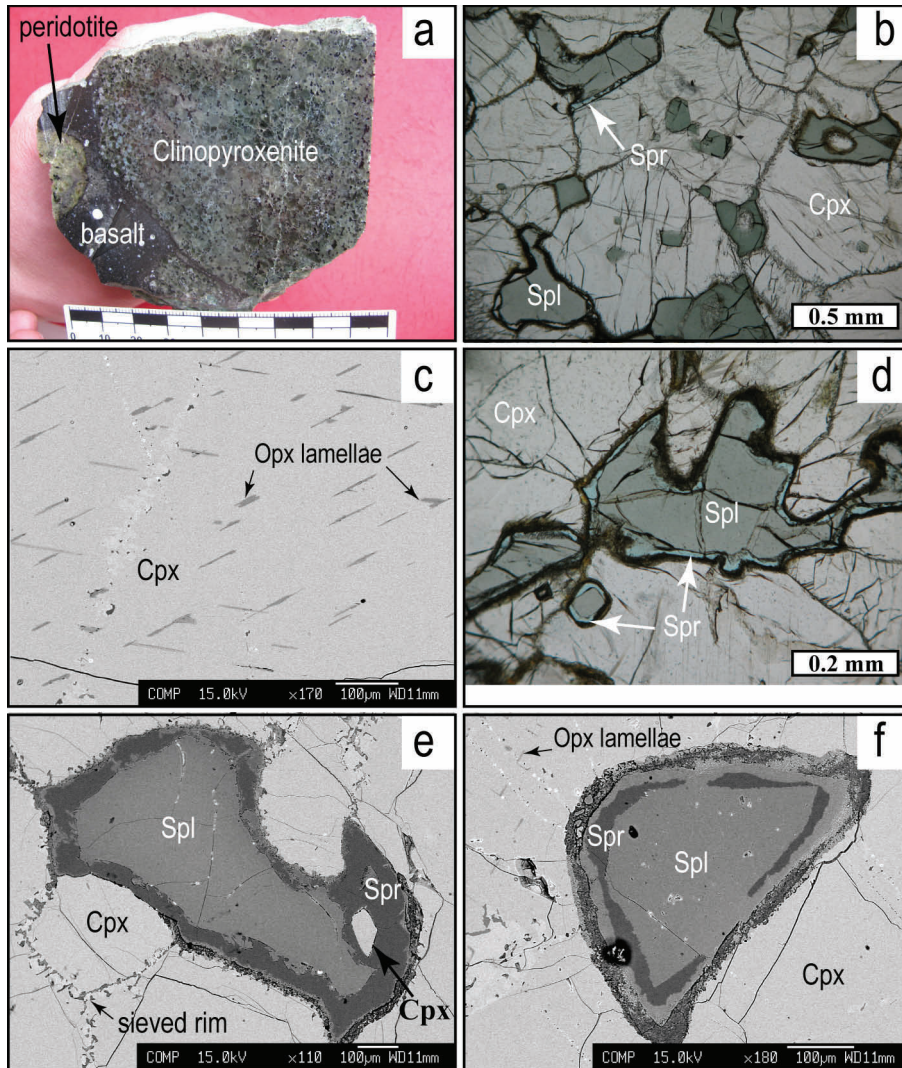
The spinel in the xenolith is magnesian and aluminous (MgO: 20.4–21.4 wt%;  $\text{Al}_2\text{O}_3$ : 64.9–66.8 wt%) with low  $\text{Cr}_2\text{O}_3$  (0.76–1.03 wt%) and  $\text{TiO}_2$  (<0.06 wt%) and moderate FeO content (9.86–12.1 wt%) with Mg no. = 75.2–79.6. The  $\text{Fe}^{3+}$  content of spinel is remarkably lower than  $\text{Fe}^{2+}$  (Table 2). Most spinel grains, including those with sapphire rims, are compositionally homogeneous, but some show Al depletion and Fe enrichment in the rims relative to the cores (Table 2).

Two sapphire grains show nearly identical concentrations of  $\text{SiO}_2$  (14.2–14.5 wt%),  $\text{Cr}_2\text{O}_3$  (0.41–0.49 wt%),  $\text{Al}_2\text{O}_3$  (60.1–62.3 wt%), MgO (18.8–19.5 wt%), FeO (3.93–4.01 wt%), MnO (0.03–0.04 wt%), CaO (0.04–0.06 wt%), and  $\text{Na}_2\text{O}$  (0.01–0.07 wt%) contents with Mg no. of 89.6–89.8 (Table 2).



**FIGURE 1.** (a) Map showing the location of the Hannuoba basalt field in the northern North China Craton. (b) Sample locality at Jieshaba in the basalt field (b) (modified after Zhou et al. 2002). A = Archean; J = Jurassic; K = Cretaceous; Q = Quaternary; E = Tertiary. (Color online.)





**FIGURE 2.** Photographs of the Hannuoba clinopyroxenite. (a) Hand specimen of clinopyroxenite associated with a peridotite xenolith showing a curved boundary with the host basalts. (b) The contact relationship between constituent minerals. Light-green spinels (Spl) occur as inclusions in clinopyroxene (Cpx), or as discrete grains mantled by sapphirine (Spr). (c) Backscattered electron image showing orthopyroxene (Opx) exsolution lamellae in Cpx. (d) Occurrence of sapphirine as prisms surrounding spinel grain. (e) Clinopyroxenes with spongy rims. Sapphirine occurs as rims on spinel and encloses fine-grained clinopyroxene. (f) Sapphirine-rimming spinel is surrounded by secondary spinel. (Color online.)

## DISCUSSION

### Occurrence and comparison with examples elsewhere in the world

Sapphirine-bearing metamorphic rocks generally have quite complex mineral assemblages and textural relationships. In garnet-sillimanite granulites (“khondalites”), sapphirine occurs in association with quartz in the matrix, as inclusions in garnet or sillimanite, or surrounded by matrix plagioclase (e.g., Santosh et al. 2007, 2009, 2012; Guo et al. 2008; Jiao and Guo 2011; Tsunogae and Santosh 2011), and less often, as small prisms enclosing spinel grains in their cores (e.g., Lal et al. 1987). In mafic granulites, sapphirine occurs with clinopyroxene, orthopyroxene, plagioclase, spinel, and garnet as major minerals. The sapphirine may form poikiloblasts, sometimes euhedral, or occur as coronae or symplectites (Arima and Barnett 1984; Christy 1989; Gregoire et al. 2001). Mutual exsolution and intergrowth of sapphirine and clinopyroxene has also been observed (Christy 1989; Lal 1997; Sutherland et al. 2003).

Among ultramafic bulk compositions, sapphirine has been observed only in pyroxenites with relatively simple mineralogy compared to their occurrence in mafic and felsic rocks. The host pyroxenites from orogenic massifs usually contain clinopyroxene, orthopyroxene, spinel, garnet, hornblende, phlogopite, corundum, and plagioclase (Kornprobst et al. 1990; Okay 1994; Morishita et al. 2001). The mineral assemblage of the sapphirine-bearing pyroxenite xenoliths reported by Meyer and Brookins (1976) and Griffin and O’Reilly (1986) consists of orthopyroxene, clinopyroxene, spinel, garnet, and plagioclase. In most cases, sapphirine is associated with, and may replace, spinel (Meyer and Brookins 1976; Griffin and O’Reilly 1986; Kornprobst et al. 1990). However, the Hannuoba clinopyroxenite reported here has a simpler mineral assemblage that consists of clinopyroxene-spinel-sapphirine; all sapphirines occur as rims on spinel (Figs. 2b–2f). The mineralogical constitution and microstructural features of the Hannuoba xenolith suggest that the sapphirine in this rock is genetically related to spinel, and formed either during the cooling/uplift process or through the reaction of spinel with a clinopyroxene solid solution.

**TABLE 1.** Bulk composition of the sapphire-bearing clinopyroxenite xenolith (JSB10-57) from Hannuoba

| SiO <sub>2</sub> | TiO <sub>2</sub> | Al <sub>2</sub> O <sub>3</sub> | TFe <sub>2</sub> O <sub>3</sub> | MnO  | MgO  | CaO  | Na <sub>2</sub> O | K <sub>2</sub> O | P <sub>2</sub> O <sub>5</sub> | LOI  | Total | FeO  | Fe <sub>2</sub> O <sub>3</sub> | Mg no. |
|------------------|------------------|--------------------------------|---------------------------------|------|------|------|-------------------|------------------|-------------------------------|------|-------|------|--------------------------------|--------|
| 40.3             | 0.11             | 20.43                          | 5.21                            | 0.10 | 15.2 | 16.2 | 0.93              | 0.13             | 0.02                          | 0.76 | 99.3  | 1.91 | 3.11                           | 85.3   |

**TABLE 2.** Chemical composition of minerals in the sapphire-bearing clinopyroxenite xenolith (JSB10-57) from Hannuoba

| Mineral<br>Position            | Spl   |        |         | Spr   |        |         | Cpx   |        |       | Opx lamella |        |       |        |        |       |       |       |
|--------------------------------|-------|--------|---------|-------|--------|---------|-------|--------|-------|-------------|--------|-------|--------|--------|-------|-------|-------|
|                                | core  | mantle | Spl rim | core  | mantle | Spl rim | core  | mantle | rim   | core        | mantle | rim   | in Spr | in Cpx |       |       |       |
| SiO <sub>2</sub>               | 0.00  | 0.03   | 14.4    | 0.00  | 0.00   | 14.2    | 0.03  | 0.00   | 0.00  | 49.7        | 49.7   | 51.0  | 49.8   | 50.0   | 50.9  | 50.6  | 52.4  |
| TiO <sub>2</sub>               | 0.06  | 0.04   | 0.02    | 0.00  | 0.03   | 0.03    | 0.00  | 0.03   | 0.00  | 0.13        | 0.14   | 0.12  | 0.13   | 0.14   | 0.16  | 0.15  | 0.02  |
| Al <sub>2</sub> O <sub>3</sub> | 66.5  | 66.0   | 62.3    | 66.2  | 66.8   | 60.1    | 66.5  | 66.0   | 64.9  | 9.86        | 10.4   | 5.94  | 10.1   | 9.76   | 5.74  | 8.67  | 7.29  |
| Cr <sub>2</sub> O <sub>3</sub> | 0.99  | 1.03   | 0.49    | 0.94  | 0.98   | 0.41    | 0.83  | 0.79   | 0.76  | 0.13        | 0.11   | 0.08  | 0.15   | 0.13   | 0.10  | 0.10  | 0.04  |
| FeO                            | 10.7  | 10.4   | 4.01    | 10.7  | 9.86   | 3.93    | 11.0  | 11.3   | 12.1  | 3.04        | 3.02   | 3.11  | 3.04   | 3.00   | 3.03  | 2.80  | 7.78  |
| MnO                            | 0.11  | 0.06   | 0.04    | 0.12  | 0.04   | 0.03    | 0.08  | 0.17   | 0.23  | 0.10        | 0.09   | 0.11  | 0.13   | 0.10   | 0.07  | 0.02  | 0.15  |
| MgO                            | 21.2  | 20.9   | 19.5    | 21.4  | 21.4   | 18.8    | 21.3  | 20.6   | 20.4  | 13.2        | 13.1   | 15.1  | 13.1   | 13.2   | 15.6  | 14.0  | 30.9  |
| CaO                            | 0.01  | 0.03   | 0.04    | 0.04  | 0.02   | 0.06    | 0.01  | 0.00   | 0.03  | 20.8        | 20.8   | 22.8  | 21.2   | 21.3   | 23.0  | 21.6  | 0.51  |
| Na <sub>2</sub> O              | 0.02  | 0.02   | 0.01    | 0.02  | 0.00   | 0.07    | 0.00  | 0.02   | 0.00  | 1.65        | 1.54   | 0.60  | 1.68   | 1.55   | 0.49  | 1.30  | 0.05  |
| K <sub>2</sub> O               | 0.00  | 0.01   | 0.01    | 0.02  | 0.00   | 0.01    | 0.00  | 0.00   | 0.00  | 0.00        | 0.00   | 0.00  | 0.01   | 0.01   | 0.01  | 0.01  | 0.00  |
| NiO                            | 0.22  | 0.24   | 0.08    | 0.21  | 0.14   | 0.07    | 0.24  | 0.19   | 0.16  | 0.00        | 0.04   | 0.00  | 0.00   | 0.09   | 0.00  | 0.03  | 0.00  |
| Total                          | 99.8  | 98.8   | 100.9   | 99.6  | 99.3   | 97.7    | 99.9  | 99.1   | 98.6  | 98.5        | 98.9   | 98.9  | 99.3   | 99.3   | 99.0  | 99.3  | 99.2  |
| O =                            | 4     | 4      | 20      | 4     | 4      | 20      | 4     | 4      | 4     | 6           | 6      | 6     | 6      | 6      | 6     | 6     | 6     |
| Si                             | 0.000 | 0.001  | 1.669   | 0.000 | 0.000  | 1.700   | 0.001 | 0.000  | 0.000 | 1.819       | 1.814  | 1.871 | 1.811  | 1.821  | 1.861 | 1.843 | 1.834 |
| Ti                             | 0.001 | 0.001  | 0.002   | 0.000 | 0.000  | 0.003   | 0.000 | 0.001  | 0.000 | 0.004       | 0.004  | 0.003 | 0.004  | 0.004  | 0.004 | 0.004 | 0.001 |
| Al                             | 1.958 | 1.964  | 8.509   | 1.952 | 1.972  | 8.478   | 1.954 | 1.964  | 1.946 | 0.425       | 0.448  | 0.257 | 0.433  | 0.419  | 0.248 | 0.372 | 0.300 |
| Cr                             | 0.020 | 0.021  | 0.045   | 0.019 | 0.019  | 0.039   | 0.016 | 0.016  | 0.015 | 0.004       | 0.003  | 0.002 | 0.004  | 0.004  | 0.003 | 0.003 | 0.001 |
| Fe <sup>3+</sup>               | 0.022 | 0.014  | 0.108   | 0.031 | 0.007  | 0.096   | 0.028 | 0.020  | 0.038 | 0.043       | 0.022  | 0.035 | 0.053  | 0.038  | 0.054 | 0.024 | 0.034 |
| Fe <sup>2+</sup>               | 0.223 | 0.219  | 0.280   | 0.223 | 0.206  | 0.297   | 0.229 | 0.238  | 0.257 | 0.050       | 0.070  | 0.060 | 0.040  | 0.054  | 0.039 | 0.062 | 0.194 |
| Mn                             | 0.002 | 0.001  | 0.004   | 0.003 | 0.001  | 0.003   | 0.002 | 0.004  | 0.005 | 0.003       | 0.003  | 0.003 | 0.004  | 0.003  | 0.002 | 0.000 | 0.005 |
| Mg                             | 0.791 | 0.787  | 3.367   | 0.798 | 0.797  | 3.352   | 0.793 | 0.773  | 0.772 | 0.719       | 0.712  | 0.827 | 0.709  | 0.716  | 0.853 | 0.758 | 1.610 |
| Ca                             | 0.000 | 0.001  | 0.005   | 0.001 | 0.000  | 0.008   | 0.000 | 0.000  | 0.001 | 0.816       | 0.814  | 0.898 | 0.824  | 0.830  | 0.900 | 0.842 | 0.019 |
| Na                             | 0.001 | 0.001  | 0.002   | 0.001 | 0.000  | 0.016   | 0.000 | 0.001  | 0.000 | 0.117       | 0.109  | 0.042 | 0.119  | 0.109  | 0.035 | 0.091 | 0.003 |
| K                              | 0.000 | 0.000  | 0.001   | 0.001 | 0.000  | 0.002   | 0.000 | 0.000  | 0.000 | 0.000       | 0.000  | 0.000 | 0.000  | 0.000  | 0.001 | 0.000 | 0.000 |
| Ni                             | 0.005 | 0.005  | 0.007   | 0.004 | 0.003  | 0.007   | 0.005 | 0.004  | 0.003 | 0.000       | 0.001  | 0.000 | 0.000  | 0.003  | 0.000 | 0.001 | 0.000 |
| Total                          | 3.022 | 3.014  | 14.000  | 3.031 | 3.007  | 14.000  | 3.028 | 3.020  | 3.038 | 4.000       | 4.000  | 4.000 | 4.000  | 4.000  | 4.000 | 4.000 | 4.000 |

Both sapphire and its coexisting minerals of different origins (crustal- and mantle-derived) show distinct variations in chemical composition. For example, the sapphire in felsic granulites (khondalites) generally contain higher SiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub>, but lower MgO and Cr<sub>2</sub>O<sub>3</sub> contents, whereas those in mafic granulites display large compositional ranges in Al<sub>2</sub>O<sub>3</sub> (62–71 wt%), MgO (16–21 wt%), and Cr<sub>2</sub>O<sub>3</sub> (0–0.5 wt%) contents (Figs. 3a–3c). Second, coexisting clinopyroxenes in mafic granulites are diopsides with high ferrosilite (Fs) end-member (>15), low Al<sub>2</sub>O<sub>3</sub> (<10 wt%) and Na<sub>2</sub>O (<7 wt%) contents, and large Mg no. variation (70–90) (Figs. 3d–3f). Third, spinels in both felsic and mafic granulites are low in Mg no., show a wide compositional range, and are relatively poor in Al<sub>2</sub>O<sub>3</sub> (Fig. 3g). In contrast, the sapphire grains in pyroxenite xenoliths have lower Al<sub>2</sub>O<sub>3</sub> (<68 wt%) but higher MgO (19–21 wt%) and Cr<sub>2</sub>O<sub>3</sub> (>0.1 wt%), and their SiO<sub>2</sub> contents are also higher at a given Al<sub>2</sub>O<sub>3</sub> level (Figs. 3a–3c). The clinopyroxenes in the pyroxenites show low Fs (<15) and high wollastonite (Wo) end-members (Fig. 3e). Their Al<sub>2</sub>O<sub>3</sub> and Na<sub>2</sub>O contents vary widely, whereas their higher Mg nos. are restricted to a relatively narrow range (Figs. 3e and 3f) indicating their mantle origin. The clinopyroxenes in the spongy rims of this study have extremely low Al<sub>2</sub>O<sub>3</sub> and Na<sub>2</sub>O contents that probably resulted from incipient melting (e.g., Griffin and O'Reilly 1986; Su et al. 2011). Finally, spinels in the pyroxenites are apparently rich in MgO and Al<sub>2</sub>O<sub>3</sub> (Fig. 3g).

In summary, crustal-derived sapphire-bearing rocks usually possess complex mineral assemblages and occurrences and display large compositional variations, consistent with their metamorphic origin during multi-stage processes at crustal depths

(Powell and Holland 1988; Christy 1989; Mattioli et al. 1996; Santosh et al. 2009, 2012). On the other hand, sapphire-bearing mantle-derived pyroxenites show relatively simple mineral assemblages and are characterized by high-Mg minerals with variable Al<sub>2</sub>O<sub>3</sub> contents.

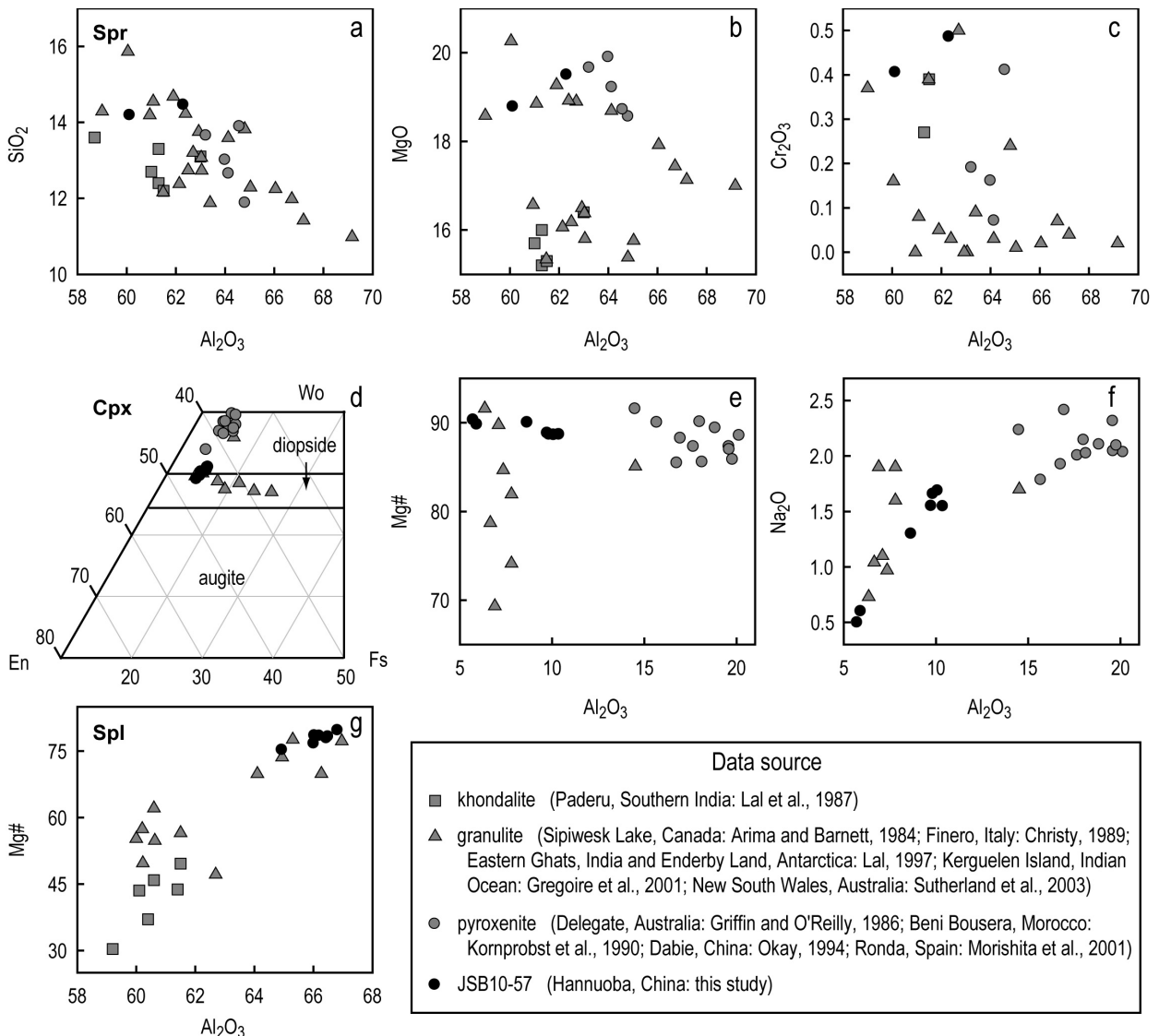
### Genesis of mantle-derived sapphire

***P-T* conditions.** The relatively simple mineralogy (Spr-Cpx-Spl; all mineral abbreviations are in the caption of Fig. 4) and the absence of appropriate mineral assemblages suitable for thermobarometry in the Hannuoba clinopyroxenite do not allow any precise *P-T* estimates for the Hannuoba clinopyroxenite. Christy's (1989) compilation from experimentally determined and calculated results define the *P-T* conditions for Spr-Cpx associations as 9–14 kbar and 800–900 °C (Fig. 4a; see references in Christy 1989). The continuous univariant spinel-sapphire transition extends from 1 kbar, around 650 °C through 10 kbar, 770 °C (Seifert 1974) to the invariant point at 22 kbar, 880 °C (Ackermann et al. 1975), which is constrained by the reaction of Clin + Crn + Spl = Spr + H<sub>2</sub>O (Fig. 4a; Okay 1994). Gasparik's study on the reaction of Grt = Spl + Opx + An ± Cpx ± Spr revealed that sapphire could be stable at 9–10 kbar over the temperature range of 700–1000 °C (Gasparik 1984). Taking into account the reaction of Grt + Fo = Cpx + Opx + Spl, a gray line-filled *P-T* field bound by 8–15 kbar and 800–900 °C is obtained for the coexisting Cpx + Spl + Spr assemblage in the present study (Fig. 4a). This field is divided into two parts by the spinel to garnet pyroxenite transition established by Herzberg (1978) at ca. 10 kbar. The lower part is the stability field for sapphire in

garnet pyroxenite, and the upper part is the field for sapphirine-bearing spinel pyroxenite. This is in good agreement with *P-T* estimates for natural rocks such as: 10–14 kbar, 800–1000 °C from the Stockdale locality (Meyer and Brookins 1976); 14 kbar, 1005 °C for Delegate (Griffin and O'Reilly 1986); and ca. 10 kbar, 800–900 °C for Ronda (Morishita et al. 2001).

Previous studies on the Hannuoba pyroxenites yielded temperatures in the range of 850–1000 °C estimated by mineral pair thermometry (e.g., Chen et al. 2001; Liu et al. 2003; Zheng et al. 2009). The spinel pyroxenite occurs at 30–40 km depths, which is regarded as the crust-mantle transition zone, and garnet pyroxenite is located below at 40–60 km depths (Fig. 4b; Chen et al. 2001; Fan et al. 2005). We thus infer that the sapphirine-bearing spinel clinopyroxenite from Hannuoba formed at about 10 kbar, 850–900 °C.

**Petrologic and geochemical constraints.** Pyroxenites cover a variety of rocks ranging from orthopyroxenites, through websterites, to clinopyroxenites, and are generally considered to have formed under high pressure in melts flowing along conduits in the mantle (e.g., Downes 2007). However, in recent years, a metasomatic origin by means of peridotite-melt interaction has received wide attention (Garrido and Bodinier 1999; Liu et al. 2005; Sobolev et al. 2005; Zhang 2009; Zhang et al. 2010). Liu et al. (2005) and Tang et al. (2011) ascribed the formation of Hannuoba pyroxenite to peridotite-melt interaction based on petrological and geochemical characteristics. In terms of mineral chemistry, clinopyroxenes in the pyroxenites display high Mg no. and large variations in some major oxides, such as SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, and CaO, resembling those in peridotite (Figs. 5a–5c). Similar features are also observed in spinels (Fig. 5d).



**FIGURE 3.** A comparison of the compositions of sapphirine, clinopyroxene, and spinel from different settings including khondalite, granulite, and pyroxenite. See text for details.



Furthermore, sample JSB10-57 shows distinct bulk  $\text{SiO}_2/\text{Al}_2\text{O}_3$  ratio relative to the Hannuoba sapphirine-free pyroxenites and host basalts (Fig. 5e; Zheng et al. 2009 and references therein), and its low  $\text{Fe}^{2+}/\text{Fe}^{3+}$  character in the bulk-rock composition is mainly controlled by spinel composition. Therefore, we consider that the pyroxenites from Hannuoba are probably genetically linked to the peridotites through reaction with an Al-Ca-Mg-rich melt. Alternatively, the oxygen ratio of sapphirine lies between those of spinel and pyroxene, and may be produced by a sliding reaction:  $2\text{MgAl}_2\text{O}_4$  (Spl) +  $2\text{Mg}_{0.75}\text{Al}_{2.5}\text{Si}_{0.75}\text{O}_6$  (in pyroxene) =  $\text{Mg}_{3.5}\text{Al}_9\text{Si}_{1.5}\text{O}_{20}$  (Spr). The unrealistic, highly aluminous pyroxene composition can be rewritten:  $2\text{MgAl}_2\text{O}_4$  (Spl) +  $2.5\text{MgAlSiO}_6$  (in Cpx or Opx) =  $\text{Mg}_{3.5}\text{Al}_9\text{Si}_{1.5}\text{O}_{20}$  (Spr) +  $0.5\text{Mg}_2\text{Si}_2\text{O}_6$  (in Cpx or Opx) (Christy 1989). Thus, sapphirine grows at the expense of spinel, and clinopyroxene loses Mg and Al, becoming closer to diopside in composition.

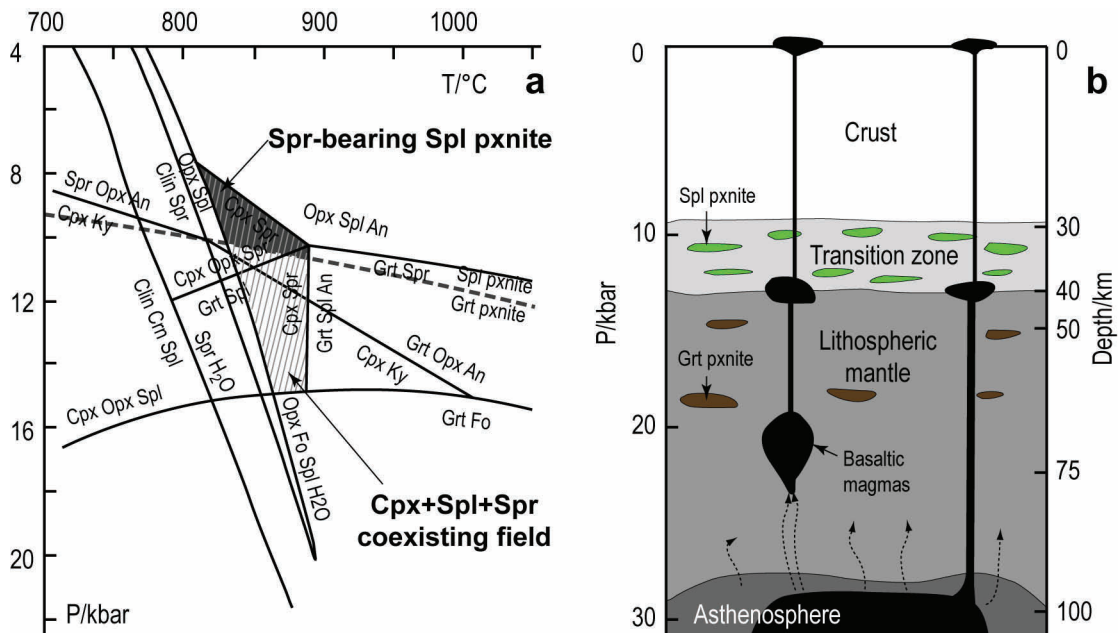
The occurrences of sapphirine in pyroxenites, including both this study and those in previous reports from elsewhere, indicate that the Spr-Cpx-Spl association developed from spinel clinopyroxenite. The available mineral-phase equilibria reactions cannot adequately explain this association without the involvement of a melt phase. This further confirms that an interaction between peridotite and melt must have taken place during pyroxenite formation. The xenolith might have undergone extensive subsolidus re-equilibration as indicated by the presence of exsolution lamellae, coronas, and intergrowth textures (Fig. 2; Mattioli et al. 1996; Gregoire et al. 2001). Furthermore, despite the abundant occurrence of garnet pyroxenites and spinel pyroxenites in the Hannuoba basalts, sapphirine-bearing rocks have not been reported until now. Therefore, only under unusual conditions can sapphirine develop in mantle depths, i.e., restricted  $P$ - $T$ , specific

bulk composition, and appropriate interaction with mantle peridotites. These features possibly explain why sapphirine is rarely found in mantle-derived rocks.

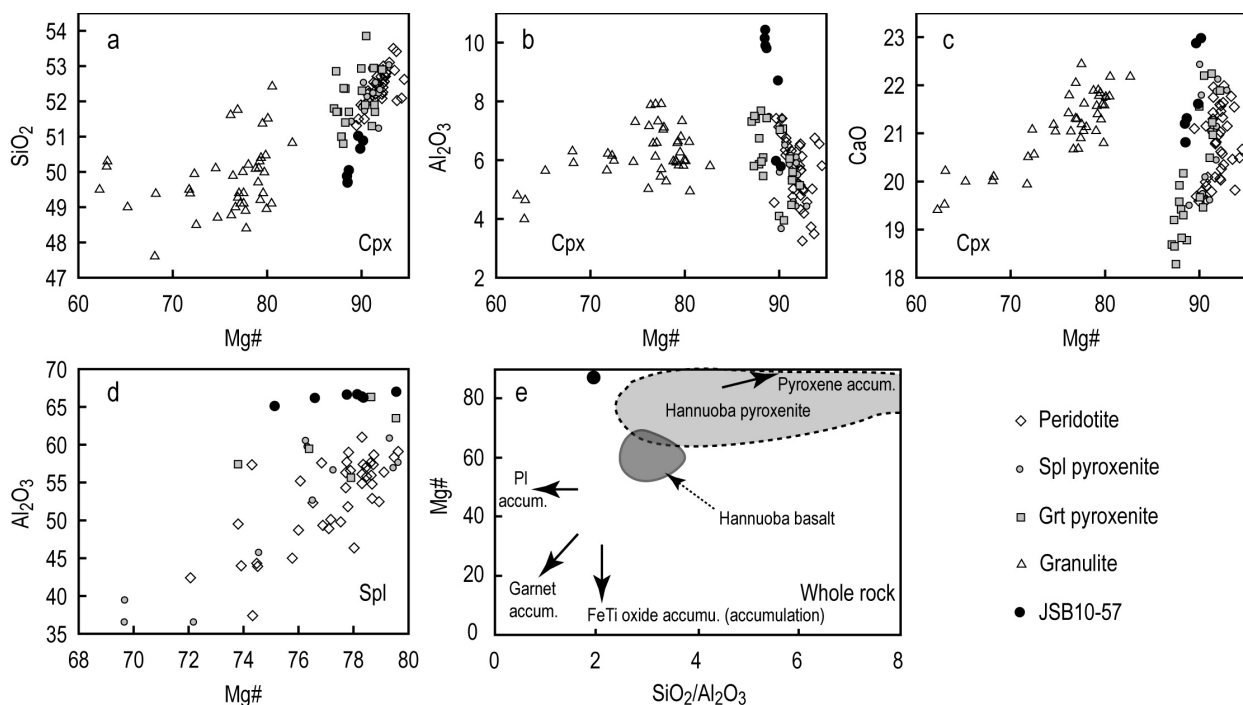
### Geological significance

The defined  $P$ - $T$  conditions (Fig. 4a) indicate that the mantle depths where the sapphirine equilibrated are typically characterized by elevated geothermal gradients and juvenile materials (Ackermann et al. 1975; Mattioli et al. 1996). This is in accordance with the presence of pyroxenites (with or without sapphirine) in the crust-mantle transition zone beneath the Hannuoba area (Fig. 4b) where the geothermal gradient is computed to be much higher than the oceanic geotherm (Chen et al. 2001; Su et al. 2007). The 10 km thick transitional zone located at about 30 to 40 km beneath the Hannuoba area (Fig. 4b) is best explained by varying degrees of mixing through voluminous basaltic magmas into the mantle low-rock. In addition, the intermediate-mafic granulite xenoliths entrained by Hannuoba basalts carrying abundant Mesozoic zircons are noted to be compositionally high in Mg no. (e.g., Zheng et al. 2009; Jiang and Guo 2010; Liu et al. 2010). These observations agree with the magma underplating model (e.g., Zhou et al. 2002; Fan et al. 2005; Liu et al. 2005; Zheng et al. 2009; Zhang et al. 2012); in addition to providing the heat input, the magmas might also have contributed to the compositional modifications required to equilibrate sapphirine in these rocks.

The high Mg, Al, and Ca contents of sapphirine-bearing mantle rocks have been evaluated in various studies. Morishita et al. (2001) suggested that such compositional characteristics result mainly from the incorporation of recycled crust. However, Gregoire et al. (1994, 1998) explained the occurrence of sapphirine



**FIGURE 4.** (a)  $P$ - $T$  diagram showing the equilibrium curves of sapphirine-related reactions compiled from Ackermann et al. (1975), Herzberg (1978), Christy (1989, and references therein), and Okay (1994). (b) Schematic diagram showing the distribution of rock-types and inferred lithospheric mantle structure beneath Hannuoba (modified from Chen et al. 2001). Spr = sapphirine; Spl = spinel; pxnite = pyroxenite; Opx = orthopyroxene; An = plagioclase; Clin = clinoclhorite; Ky = kyanite; Grt = garnet; Crn = corundum; Fo = forsterite. (Color online.)



**FIGURE 5.** (a–c) Comparison of the compositions of clinopyroxene from the Hannuoba peridotite, spinel pyroxenite, garnet pyroxenite, granulite, and sapphirine-bearing spinel clinopyroxenite (JSB10-57). Data are compiled from Chen et al. (2001), Liu et al. (2003), Rudnick et al. (2004), Fan et al. (2005), and Tang et al. (2007). (d) Comparison of the compositions of spinel from the Hannuoba peridotite, spinel pyroxenite, garnet pyroxenite, and sapphirine-bearing spinel clinopyroxenite (JSB10-57). Data are compiled from Chen et al. (2001), Rudnick et al. (2004), and Tang et al. (2007). (e) Mg no. vs.  $\text{SiO}_2/\text{Al}_2\text{O}_3$  diagram for Hannuoba pyroxenite xenoliths. The fields of Hannuoba pyroxenite and basalt are from Zheng et al. (2009, and references therein). Note that Mg no. is defined as  $\text{Mg}/(\text{Mg}+0.85\text{Fe}_{\text{total}})$ , to be consistent with Kempton and Harmon (1992). Unbroken arrows show the generalized direction of compositional change with accumulation of the mineral phase indicated.

from oceanic components and Ackermann et al. (1975) inferred that water also plays a major role in producing the sapphirine (Fig. 4a). Subducted oceanic crust has also been proposed as the source of parental magmas for pyroxenites to account for their compositional diversity (Medaris et al. 1995; Xu 2002; Liu et al. 2010; Yu et al. 2010; Zhang et al. 2010).

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